Gravitinos, Reheating and the Matter-Antimatter Asymmetry of the Universe

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OUTLINE

- THE MATTER-ANTIMATTER ASYMMETRY OF THE UNIVERSE
- INFLATION AND REHEATING
- THE GRAVITINO PROBLEM, AND T_{reh}

REHEATING, GRAVITINOS AND THE M-A ASYMMETRY

- A WAY OUT: DETAILED VIEW OF REHEATING
- ANOTHER WAY OUT: DELAYED THERMALISATION
- GRAVITINO PROBLEM AGAIN

CONCLUSION

PREAMBLE

A BRIEF HISTORY OF OUR UNIVERSE

OBSERVATIONS + GENERAL THEORY OF RELATIVITY

14 b yr, COMPOSITION, EXPANDING, PAST – HOT AND DENSE

A BRIEF HISTORY OF OUR UNVIERSE

- First second hot primordial plasma of electrons, . photons, quarks/protons, neutrons, dark matter, ...
- 1 s 3 min light nuclei (helium, lithium, ..)
- 400,000 years Atoms form, CMBR
- 300 million years First stars form
- 1 billion years First galaxies form
- 9 billion years Solar system formed
- 14 billion years Today

THE FIRST SECOND

• 10^{-44} s – Planck time (E ~ 10^{19} GeV) [Q Gravity]

Grand Unified Theory

 10⁻³⁸ s – GUT Phase Transition (E ~ 10¹⁶ GeV, T ~ 10²⁹ K)

Standard Model [q, I, H, GB] /Modified SM

- 10⁻¹¹ s Electroweak Phase Transition (E ~ 100 GeV, T ~ 10¹⁵ K)
- $10^{-6} \text{ s} \text{quarks} \rightarrow \text{protons}$, neutrons (E ~ 1 GeV, T ~ 10^{13} K)
- 1 s Primordial Nucleosynthesis begins (E~ 1 MeV, T~ 10¹⁰ K)

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MATTER-ANTIMATTER ASYMMETRY OF THE UNIVERSE

- SOLAR SYSTEM PROBES, INTERACTION OF SOLAR WIND WITH PLANETS
- MILKY WAY COSMIC RAYS
- CLUSTER (20 Mpc) GALACTIC COLLISIONS (1 Mpc = 3 x 10⁶ lt-yr) INTERGALACTIC HOT PLASMA
- UP TO 1000 Mpc COSMIC DIFFUSE GAMMA RAY SPECTRUM
 (ANNIHILATIONS AT BOUNDARY FROM z=1000 TO 20 – 380,000 YR TO 100 MILLION YR)
 (Cohen, de Rujula, Glashow) 7

MATTER-ANTIMATTER ASYMMETRY OF THE UNIVERSE

- ANTIMATTER RULED OUT TILL d~1000 Mpc
- SIZE OF OBSERVABLE UNIVERSE ~ 14000 Mpc

 $(1 \text{ Mpc} = 3 \times 10^{19} \text{ km} = 3 \times 10^{6} \text{ It-yr})$

MATTER-ANTIMATTER ASYMMETRY OF THE UNIV

- EARLY TIMES (t << 1 s = PRIM. NUCL.) EQUAL AMOUNTS OF MATTER AND ANTIMATTER
- WHERE DID THE ANTIMATTER GO? WHY THIS ASYMMETRY TODAY?
- DISEQUILIBRIUM IN THE EARLY UNIVERSE 100 M + 100 $A \rightarrow$ 103 M + 101 $A \rightarrow$ 2 M

 $X \to M$ $X \to A$

 $\rm r_M > \rm r_A$, GET MORE MATTER THAN ANTIMATTER

- X = GUT (GRAND UNIFIED THEORY) BOSONS – GUT BARYOGENESIS MASS ($M_x \sim 10^{16}$ GeV)
- X = HEAVY NEUTRINOS - LEPTOGENESIS MODELS MASS ($M_N \sim 10^{10}$ GeV)

MASS EXPRESSED AS MASS ENERGY M $\rm c^2$

1 GeV = PROTON MASS

BEYOND STANDARD MODEL PARTICLES

WHEREFROM

- GUT BOSONS ($M_X \sim 10^{16} \text{ GeV}$)
- HEAVY NEUTRINOS ($M_N \sim 10^{10} \text{ GeV}$) ?

1 GeV = PROTON MASS

WHEREFROM

- GUT BOSONS ($M_X \sim 10^{16} \text{ GeV}$)
- HEAVY NEUTRINOS $(M_N \sim 10^{10} \text{ GeV})$?

1 GeV = PROTON MASS

In the hot early Universe when temperatures were very high $(k_B T > M)$ $(k_B = 1)$

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INFLATION and **REHEATING**

INFLATION – PERIOD OF ACCELERATED EXPANSION IN THE EARLY UNIVERSE (t ~ 10^{-38} s or later)

ASSOCIATED WITH THE DYNAMICS OF A SLOWLY VARYING FIELD CALLED THE INFLATON Φ



ENERGY DENSITY DOMINATES, DETERMINES EVOL OF UNIV



DURING INFLATION, $R \sim exp(H t)$ [R IS THE SCALE FACTOR,

In expanding Univ d ~ d₁ R(t)]

n OF ALL SPECIES $\rightarrow 0$

INFLATON DECAY PRODUCTS THERMALISE, T_{reh} THERMAL BATH HAS q, I, H, dm, BSM INCLUDING GUT PARTICLES AND HEAVY NEUTRINOS REHEATING

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GRAVITINOS

 $\tilde{G} =$ SUPERSYMMETRIC PARTNER OF THE GRAVITON

SUPERSYMMETRY

- EXTENSION OF THE STANDARD MODEL (GAUGE HIERARCHY)
- SUPERPARTNERS: FERMION BOSON

PHOTON – PHOTINO, ELECTRON – SELECTRON (EQUAL m, IF SUSY)

LOCAL (spacetime dep) SUPERSYMMETRY: SUPERGRAVITY GRAVITON – GRAVITINO (\tilde{G})

BROKEN $(m_{\tilde{G}} : eV - TeV)$ 17

GRAVITINOS

 $\tilde{G} =$ SUPERSYMMETRIC PARTNER OF THE GRAVITON

PRODUCED AFTER INFLATION $t \sim 10^{-38} \,\mathrm{s} \,(m_{\tilde{G}} : \mathrm{eV} - \mathrm{TeV})$

COSMOLOGICAL CONSEQUENCES (m, n)

- STABLE : AFFECTS EXPANSION RATE, $\rho_{\tilde{G}} > \rho_c$ (L/H)
- UNSTABLE : AFFECT EXPANSION RATE PRIOR TO DECAY

DECAY PRODUCTS $\rho > \rho_c$

DESTROY LIGHT ELEMENTS ${}^{4}He$, ${}^{3}He$, D (NUCLEOSYNTHESIS)

GRAVITINO PROBLEM(S)

GRAVITINOS

 $\tilde{G} =$ SUPERSYMMETRIC PARTNER OF THE GRAVITON

PRODUCED AFTER INFLATION $t \sim 10^{-34} \,\mathrm{s} \,(m_{\tilde{G}} : \mathrm{eV} - \mathrm{TeV})$

COSMOLOGICAL CONSEQUENCES (m, n)

- STABLE : AFFECTS EXPANSION RATE, $\rho_{\tilde{G}} > \rho_c$ (L/H)
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DESTROY LIGHT ELEMENTS ${}^{4}He$, ${}^{3}He$, D (NUCLEOSYNTHESIS)

GRAVITINO PROBLEM(S) => UPPER BOUND ON $ho_{ ilde{G}} \propto n_{ ilde{G}}$

STANDARD PICTURE OF GRAVITINO PRODUCTION



INFLATION \rightarrow REHEATING (OSC. + DECAY) (T_{reh})

→ RADIATION DOMINATED UNIV (Relativistic particles)

THERMAL SCATTERING $\rightarrow G$ (gluons, quarks, squarks, gluinos) ²⁰

STANDARD CALC OF GRAVITINO PRODUCTION

CALCULATE GRAVITINO PRODUCTION IN THE RAD DOM ERA

MAINLY PRODUCED AT THE BEGINNING OF THE RAD DOM ERA WHEN $~T\sim T_{\rm reh}$, and $~n_{\tilde{G}}\propto T_{\rm reh}.$

UPPER BOUND ON $n_{\tilde{G}}$

 \Rightarrow UPPER BOUND ON T_{reb} OF 10^{6—9} GeV (MASS 100 GeV – 10 TeV)

 $k_{\rm B}$ T in GeV $k_{\rm B}$ =1 1 GeV =10¹³ K

 THE UPPER BOUND ON THE REHEAT TEMPERATURE 10⁶⁻⁹ GeV TO SUPPRESS GRAVITINO PRODUCTION

1 GeV =10¹³ K

- THE UPPER BOUND ON THE REHEAT TEMPERATURE 10⁶⁻⁹ GeV TO SUPPRESS GRAVITINO PRODUCTION
- MATTER-ANTIMATTER ASYMMETRY GENESIS MODELS REQUIRE HEAVY X, MASS 10¹⁰, 10¹⁶ GeV

1 GeV = PROTON MASS

- THE UPPER BOUND ON THE REHEAT TEMPERATURE 10⁶⁻⁹ GeV TO SUPPRESS GRAVITINO PRODUCTION
- MATTER-ANTIMATTER ASYMMETRY GENESIS MODELS REQUIRE HEAVY X, MASS 10¹⁰, 10¹⁶ GeV

DIFFICULT TO HAVE ENOUGH HEAVY X IN THE RADIATION DOMINATED UNIV AFTER REHEATING

$$n_X \sim exp(-m c^2/k_BT)$$

- THE UPPER BOUND ON THE REHEAT TEMPERATURE 10⁶⁻⁹ GeV TO SUPPRESS GRAVITINO PRODUCTION
- MATTER-ANTIMATTER ASYMMETRY GENESIS MODELS REQUIRE HEAVY X, MASS 10¹⁰, 10¹⁶ GeV

DIFFICULT TO HAVE ENOUGH HEAVY X IN THE RADIATION DOMINATED UNIV AFTER REHEATING

LOW REHEAT TEMPERATURE IS A PROBLEM FOR GUT BARYOGENESIS AND LEPTOGENESIS ²⁶

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WE FOCUS ON LEPTOGENESIS MODELS – OUT OF EQM DECAY OF *N*.

POPULAR – RELATED TO LIGHT NEUTRINO MASSES

MASS $M_N \sim 10^{10} \text{ GeV}$



TWO SPECIES NEUTRINOS AND GRAVITINOS BOTH CREATED IN THE SAME THERMAL ENVIRONMENT -- RADIATION DOMINATED UNIVERSE AFTER REHEATING WANT N (M-A ASYMMETRY) BUT NOT \tilde{G} (DECAY)

SOLUTIONS

INCREASE N

DETAILED VIEW OF REHEATING

DECREASE \tilde{G}

DELAYED THERMALISATION DURING REHEATING DUE TO SUSY FLAT DIRECTIONS

SOLUTIONS

INCREASE N

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NEW GRAVITINO PROBLEM

INCREASE \tilde{G} DUE TO SUSY FLAT DIRECTIONS 31

SOLUTION 1

INCREASE N

DETAILED VIEW OF REHEATING

NEUTRINO PRODUCTION DURING REHEATING

STANDARD CALC OF PRODUCTION ASSUMES INSTANTANEOUS INFLATON DECAY AND REHEATING.

 $T \to T_{max} \to T_{reh}$

 $T_{\rm reh}$ is the final temperature at the end of Reheating

 $T_{\rm max}\,$ CAN BE AS HIGH AS $\,$ 1000 $\,T_{\rm reh}\,$. CAN BE USED TO CREATE ENOUGH NEUTRINOS

CHUNG ET AL, DELEPINE AND SARKAR, GIUDICE ET AL

GRAVITINO PRODUCTION DURING REHEATING

STANDARD CALC OF PRODUCTION ASSUMES INSTANTANEOUS INFLATON DECAY AND REHEATING.

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IF A LARGE T_{max} CAN ENHANCE NEUTRINO PRODUCTION, CAN IT ALSO ENHANCE GRAVITINO PRODUCTION ?

GRAVITINO PRODUCTION DURING REHEATING

SOLVED THE INTEGRATED BOLTZMANN EQUATION FOR GRAVITINO PRODUCTION DURING REHEATING

$$\frac{dn_{\tilde{G}}}{dt} = -3Hn_{\tilde{G}} + \langle \Sigma_{\text{tot}} | v | \rangle n^2$$

e.g. $q + \bar{\tilde{q}} \to g + \tilde{G}$ $q + \bar{q} \to \tilde{g} + \tilde{G}$ $\tilde{q} + \bar{\tilde{q}} \to \tilde{g} + \tilde{G}$

 $q - \tilde{q}, g - \tilde{g}$ Superpartners

RESULTS

SOLVED THE INTEGRATED BOLTZMANN EQUATION FOR GRAVITINO PRODUCTION DURING REHEATING

$$\frac{dn_{\tilde{G}}}{dt} = -3Hn_{\tilde{G}} + \langle \Sigma_{\rm tot} | v | \rangle n^2$$

e.g.
$$q + \bar{\tilde{q}} \to g + \tilde{G} \qquad q + \bar{q} \to \tilde{g} + \tilde{G} \qquad \tilde{q} + \bar{\tilde{q}} \to \tilde{g} + \tilde{G}$$

DEPENDENCE ON T_{max} CANCELS OUT [UNEXPECTED]

ABUNDANCE GENERATED IS LARGE, BUT LESS THAN THE COSMOLOGICAL BOUND ON THE GRAVITINO ABUNDANCE

SOLUTION IS VIABLE

RR, SAHU

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NEW SCENARIO OF GRAVITINO PRODUCTION

IN THE PRESENCE OF SUSY FLAT DIRECTIONS

IN SOME CASES, SUPPRESS PRODUCTION

IN OTHER CASES, EXCESSIVE PRODUCTION

SUSY FLAT DIRECTIONS

STANDARD MODEL , H SCALAR (SPIN 0) MINIMISE V, $\langle H \rangle \neq 0 \Rightarrow q$, I, W, Z GET MASS HIGGS MECHANISM

SCALAR POTENTIAL V IN SUSY IS A FUNCTION OF $(H_u, H_d, \tilde{q}_i, \tilde{l}_i)$

DIRECTIONS IN FIELD SPACE OF SCALARS ALONG WHICH THE SCALAR POTENTIAL IS MINIMISED

V' = 0, POTENTIAL IS FLAT — FLAT DIRECTIONS

[POTENTIAL IS CONSTANT AND ZERO ALONG FLAT DIRECTION]



Any point on this line minimises the potential – parametrised by ψ . Note that each point corresponds to a different vacuum

SUSY FLAT DIRECTIONS

FLAT DIRECTION CORRESPONDING TO

$$\widetilde{q}$$
 = ψ , \widetilde{l} = ψ phases

REPRESENTED BY A COMPLEX SCALAR FIELD $\,\psi\,$ (AFFLECK-DINE FIELD)



FLAT DIRECTION \rightarrow QUADRATIC POT WITH CURV m₀

 $\psi_0 \neq 0~$ due to quantum fluctuations during inflation; other reasons

WHEN t_U~ t_F (or H ~ m₀), ψ oscillates, $\psi \sim 1/R^{3/2}$ Then it decays (before ewsb t~10⁻¹¹ s)

SOLUTION 2

DECREASE \tilde{G}

DELAYED THERMALISATION DURING REHEATING DUE TO SUSY FLAT DIRECTIONS

COSMOLOGICAL CONSEQUENCES

- NON-ZERO VALUE OF ψ GIVES MASS TO GAUGE BOSONS (BREAKS GAUGE SYMMETRY),
- e.g., $L \supset \tilde{q}^* \tilde{q} A A$

FLAT DIRECTION EXPECTATION VALUE CAN BE 10¹³ GEV OR HIGHER

THERMALISATION DUE TO PROCESSES MEDIATED BY GAUGE BOSONS – PHOTONS (EM), GLUONS (STRONG)

COSMOLOGICAL CONSEQUENCES

NON-ZERO VALUE OF ψ GIVES MASS TO GAUGE BOSONS (BREAKS GAUGE SYMMETRY),

e.g., $L \supset \tilde{q}^* \tilde{q} A A$

IF ALL GAUGE BOSONS GET MASS [LLddd, QuQue], IT SLOWS DOWN THERMALISATION AFTER INFLATION, LEADING TO A DILUTE PLASMA.

SUPPRESSES GRAVITINO PRODUCTION

ALLAHVERDI AND MAZUMDAR; RR AND A. SARKAR

COSMOLOGICAL CONSEQUENCES

STANDARD PICTURE OF REHEATING: INFLATON DECAYS $\rightarrow n_0 \rightarrow$ THERMALISE KINETIC EQM n_0 CHEMICAL EQM n_1 [10⁴]

FLAT DIRECTIONS: INFLATON DECAYS $\rightarrow n_0 \rightarrow$ delayed thermalisation $n \sim n_0 \ll n_1$

DILUTE PLASMA

GRAVITINOS PRODUCED BY SCATTERING OF INFLATON DECAY PRODUCTS [n.n]

$$n_{\tilde{G}} \downarrow \downarrow$$
 46

EARLIER INFLATON DECAYS AND DECAY PRODUCTS THERMALISE QUICKLY

$$q + \bar{\tilde{q}} \to g + \tilde{G} \qquad q + \bar{q} \to \tilde{g} + \tilde{G} \qquad \tilde{q} + \bar{\tilde{q}} \to \tilde{g} + \tilde{G}$$

$$\dot{n}_{\tilde{G}} = -3Hn_{\tilde{G}} + \langle \Sigma_{\text{tot}} | v | \rangle n^2 \qquad n \sim T^3$$

NOW,
$$\dot{n}_{\tilde{G}} = -3Hn_{\tilde{G}} + \int d\Pi_1 \ d\Pi_2 \ f_1 \ f_2 \ W_{12}(s)$$

 $W_{12}(s) \propto \sigma_{CM}$

$f_{1,2}$ particle distribution functions for incoming particles

RESULTS

APPROPRIATE $f_{1,2}$

SUPPRESSED GRAVITINO PRODUCTION DUE TO

A) DILUTE PLASMAB) PHASE SPACE SUPPRESSION

$$q + \bar{\tilde{q}} o g + \tilde{G} \qquad q + \bar{q} o \tilde{g} + \tilde{G} \qquad \tilde{q} + \bar{\tilde{q}} o \tilde{g} + \tilde{G}$$

OUTGOING GLUON/GLUINO HEAVY GRAVITINO PRODUCTION SHUTS OFF WHEN THE ENERGY OF INCOMING QUARKS/SQUARKS < $m_{g,\tilde{g}}_{_{48}}$

RESULTS

SUPPRESSED GRAVITINO PRODUCTION

$$Y_{\tilde{G}} = 4 \times 10^{-18}, 10^{-20} < 10^{-14}$$

COMPLETE SHUT OFF

[RR, A. SARKAR]

N↓ BUT SUFFICIENT

GRAVITINO PRODUCTION

• DETAILED VIEW OF REHEATING N^ \uparrow BUT \tilde{G} \uparrow

• DELAYED THERMALISATION IN THE PRESENCE OF SUSY FLAT DIRECTIONS N \downarrow BUT $\tilde{G} \downarrow \downarrow$

ALTERNATE SCENARIO WITH SUSY FLAT DIRECTIONS

GRAVITINO OVER-PRODUCTION

ALTERNATE SCENARIO

- IF FLAT DIRECTION EV DOES NOT BREAK ALL GAUGE SYMMETRIES, THERMALISATION WILL OCCUR
- CONSIDER A SCENARIO WITH $H_u H_d$ FLAT DIRECTION. SU(3)_C x SU(2)_L x U(1)_Y \rightarrow SU(3)_C x U(1)_{EM}
- GLUON AND GLUINO LIGHT (m ~ gT, REL), THERMAL DISTRIBUTION
- QUARK AND SQUARK HEAVY (NR), $m \approx h\psi$, ψ > 10¹³ GeV $m_{\tilde{a}}^2 - m_a^2 = m_S^2$ $m_S^2 \sim T^2 \ll m_{a,\tilde{a}}^2$ 52



- BREIT-WIGNER RESONANCE WHEN Incoming energy = E_{gluino} + E_q ≈ m_{sq}
- CROSS SECTION ~ 1 ($s-m_{sq}^2$)² + $m_{sq}^2 \Gamma^2$

 $s^{1/2}$ = E_{gluino} + $E_{q}\,$, $\,\Gamma$ = squark decay rate

GRAVITINO PROBLEM AGAIN!

- GRAVITINO ABUNDANCE GENERATED IS VERY
 LARGE AND GREATER THAN THE COSMOLOGICAL
 UPPER BOUND FOR MOST PARAMETER SPACE
- COSMOLOGICAL UPPER BOUND IS Y < 10⁻¹⁴
- FOR DIFFERENT SETS OF PARAMETERS

 $Y = 10^{-8} - 10^{-2}$

GRAVITINO PROBLEM AGAIN!

- LARGE VALUES FOR SUSY FLAT DIRECTIONS IS GENERIC. EXACERBATED GRAVITINO PROBLEM
- HAVE TO INVOKE EARLY DECAY OF FLAT
 DIRECTIONS TO AVOID CONFLICT

[MAHAJAN, RR, A. SARKAR]

CONCLUSION

- 1. POPULAR MODELS OF GENERATING THE MATTER-ANTIMATTER ASYMMETRY OF THE UNIVERSE REQUIRE A LARGE REHEAT TEMPERATURE AFTER INFLATION
- 2. BUT THAT GENERATES TOO MANY GRAVITINOS IN THE UNIVERSE
- 3. COSMOLOGISTS ARE LOOKING FOR MECHANISMS TO ENHANCE NEUTRINO ABUNDANCE/SUPPRESS GRAVITINO ABUNDANCE

CONCLUSION

- NEUTRINOS GENERATED DURING REHEATING ~ GRAVITINO ABUNDANCE GENERATED NOT TOO LARGE
- 5. GRAVITINO ABUNDANCE GENERATED IN A NON-THERMAL UNIVERSE IN THE PRESENCE OF FLAT DIRECTIONS IS SUPPRESSED
- GRAVITINO ABUNDANCE IN A THERMAL UNIVERSE WITH FLAT DIRECTIONS CAN BE LARGE – NEW SOURCE OF THE GRAVITINO PROBLEM

(DETAILS OF THE SUSY MODEL) 57

ADJUST THE REHEAT TEMP?

- GRAVITINO ABUNDANCE DECREASES BY INCREASING $\mathsf{T}_{\mathsf{REH}}$
- STANDARD PRODUCTION GRAVITINO ABUNDANCE INCREASES WITH T_{REH}