

# GRAVITATIONAL LENSING OF GRAVITATIONAL WAVES

**Parameswaran Ajith**

International Centre for Theoretical Sciences, TIFR, Bangalore

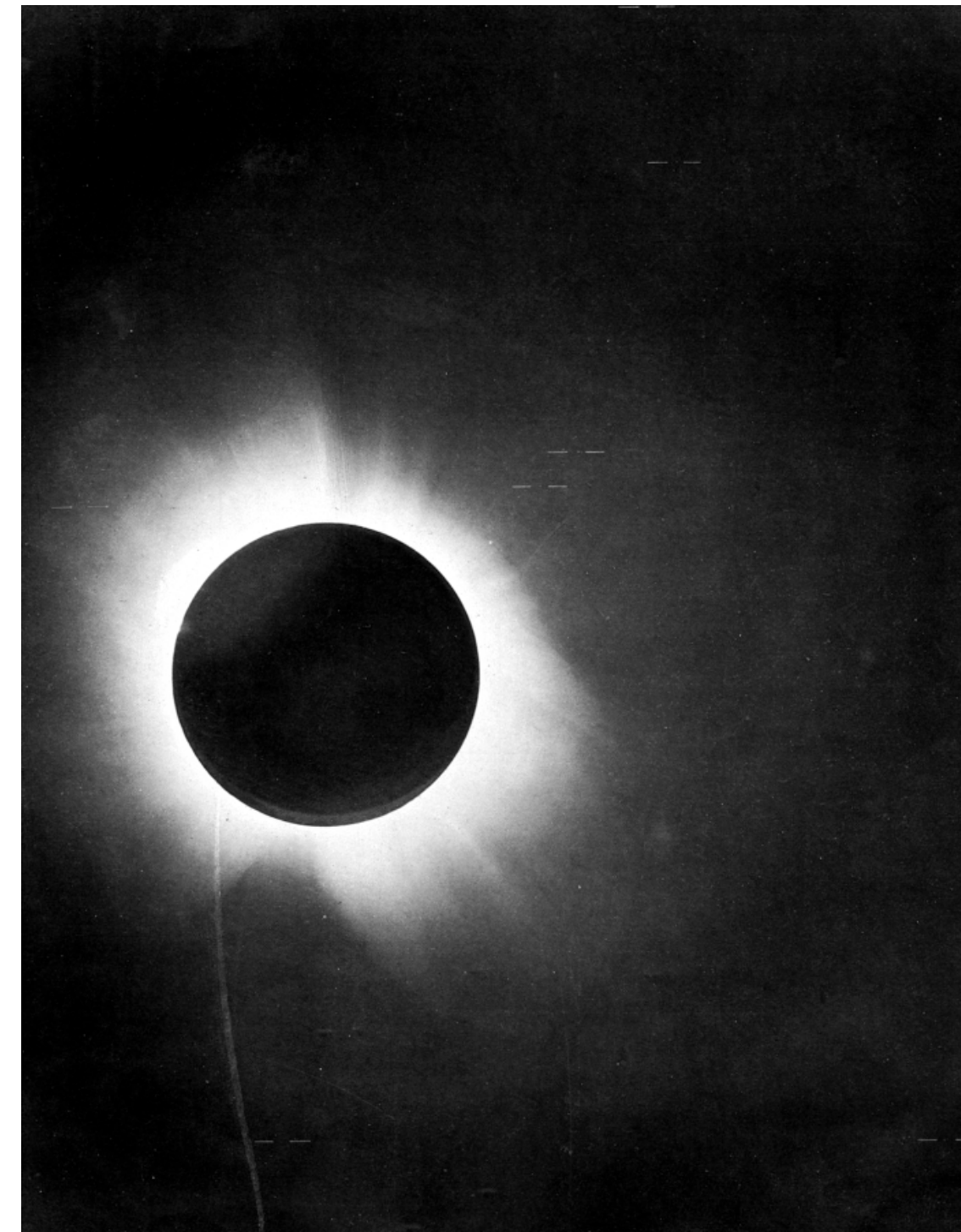
Chennai Symposium on Gravitation and Cosmology | IIT Madras | 23 Jan 2020

# 1919: Eddington's observation of gravitational light bending

---

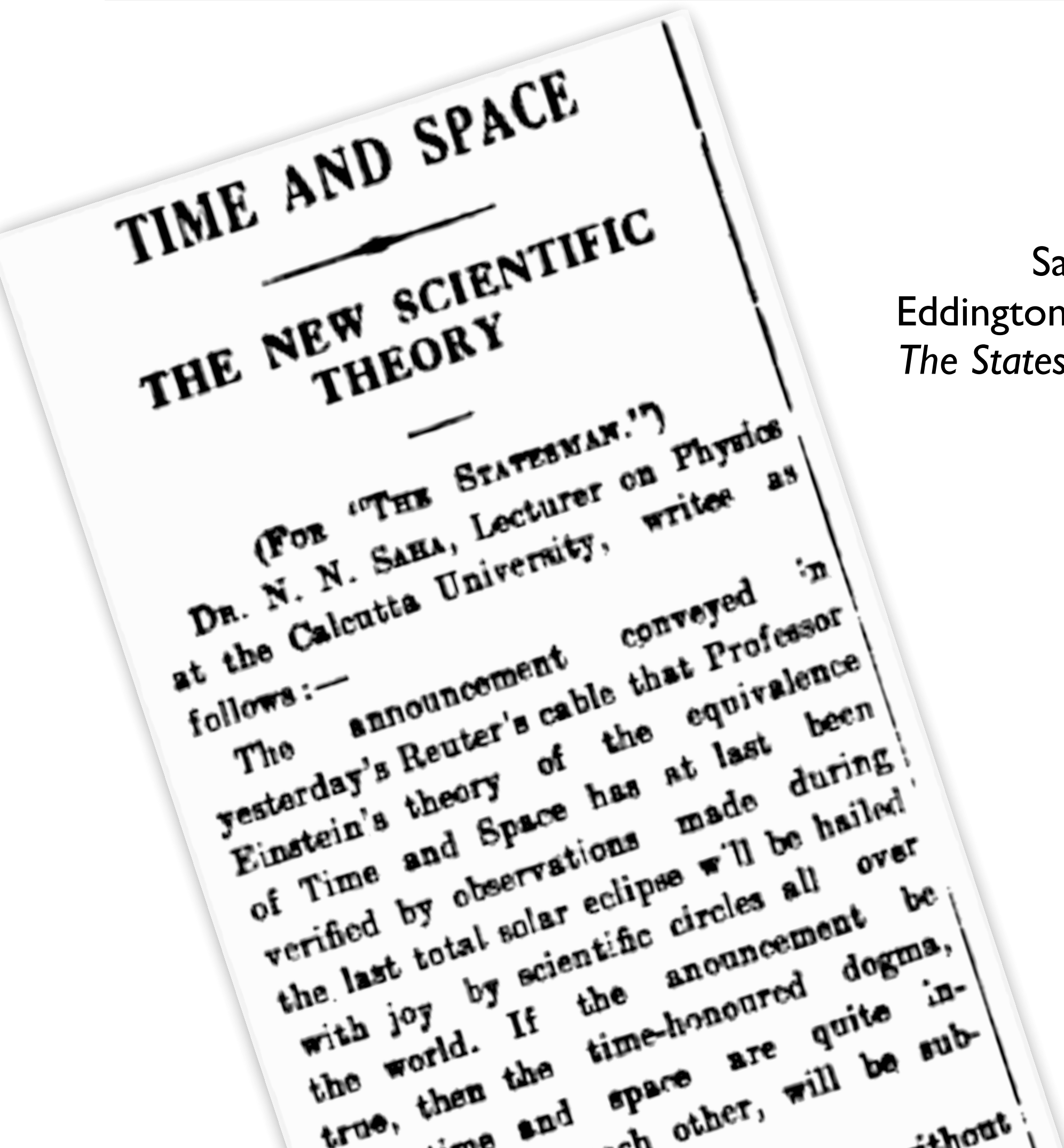


Arthur Eddington



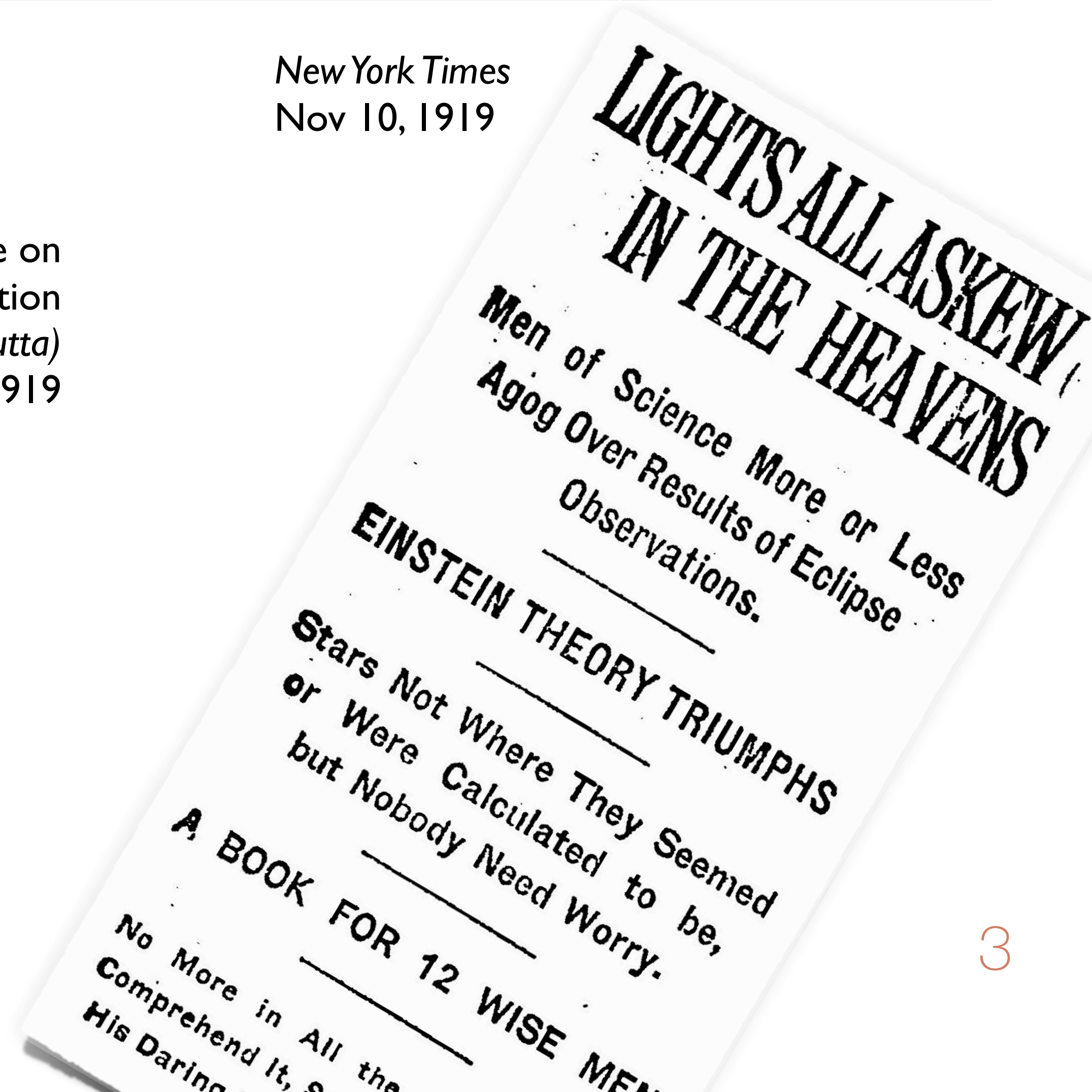
Eddington's photograph of the total solar eclipse of 29 May 1919

# Triumph of Einstein's theory



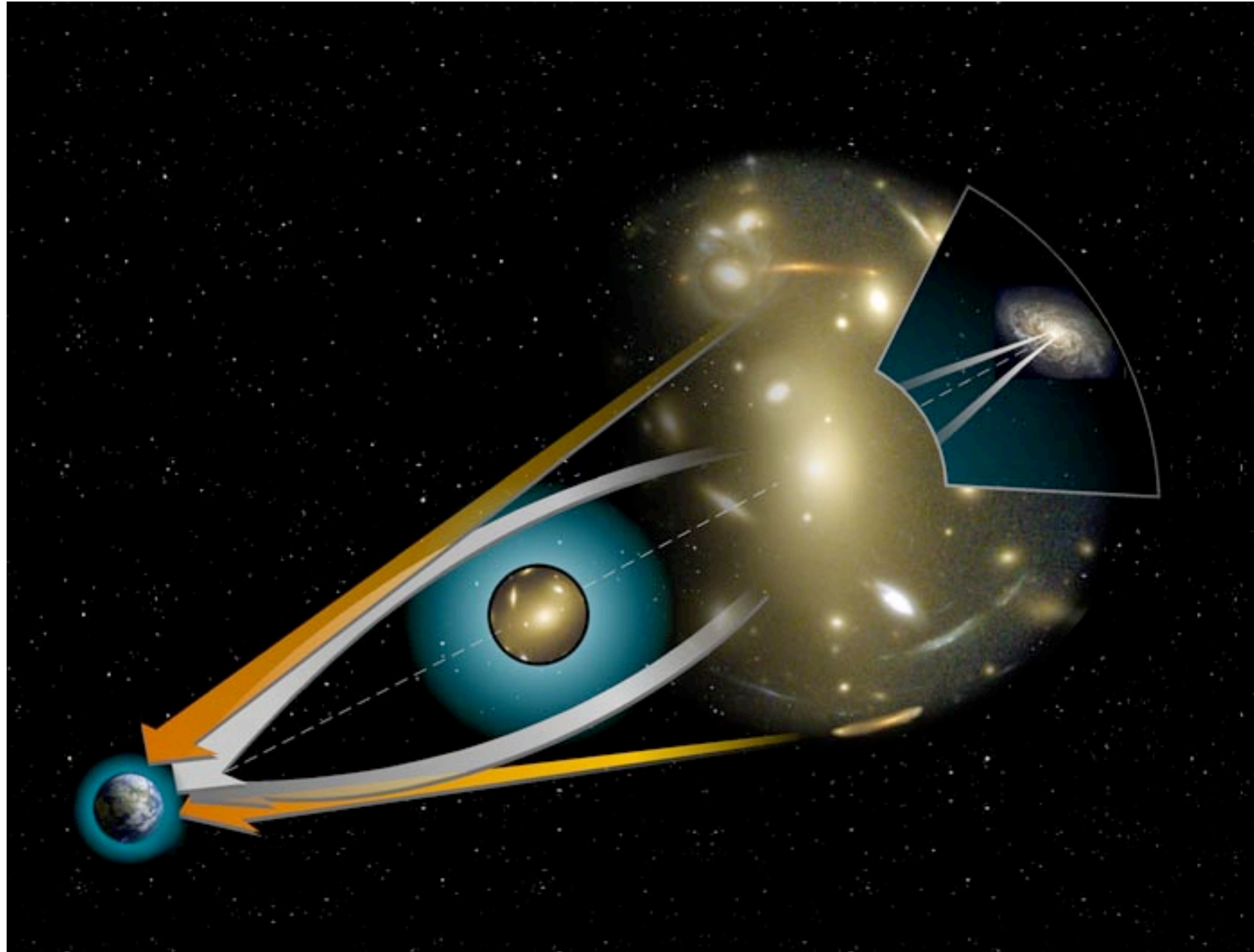
Saha's article on Eddington's observation  
*The Statesman (Calcutta)*  
13 Nov 1919

New York Times  
Nov 10, 1919



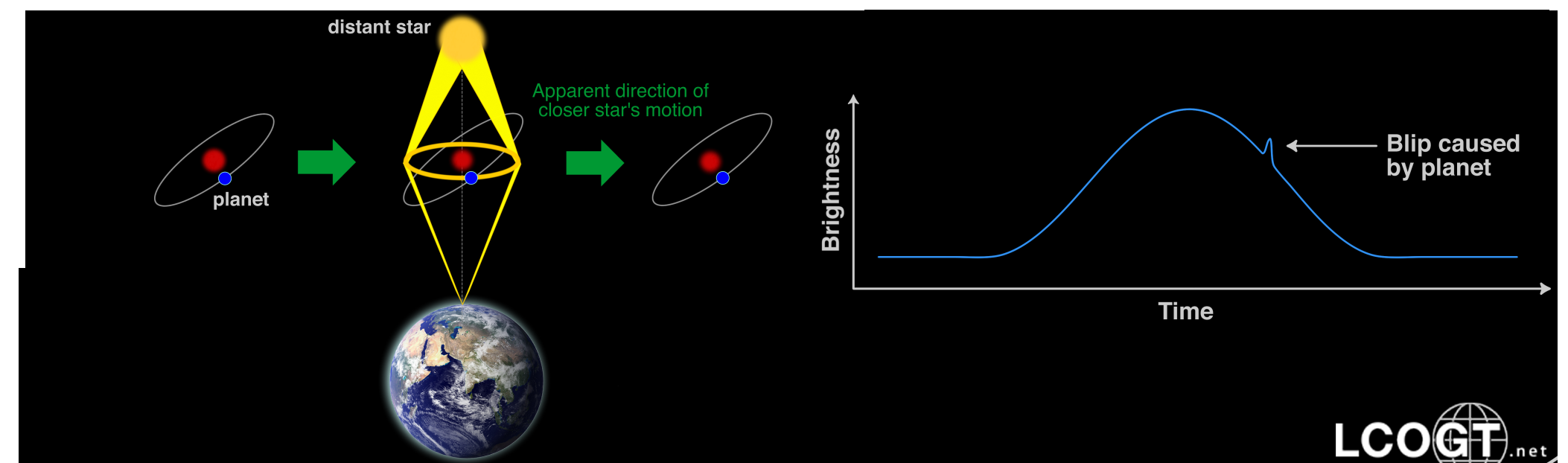
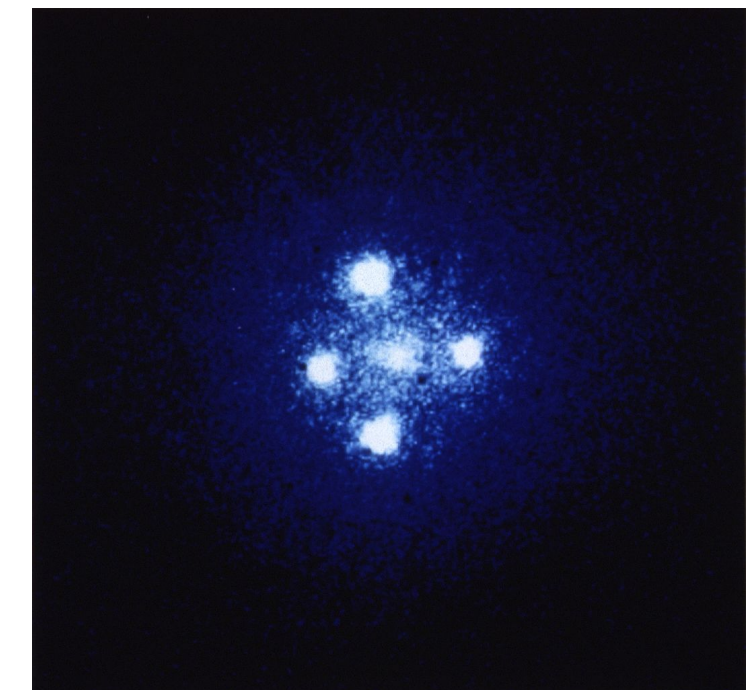
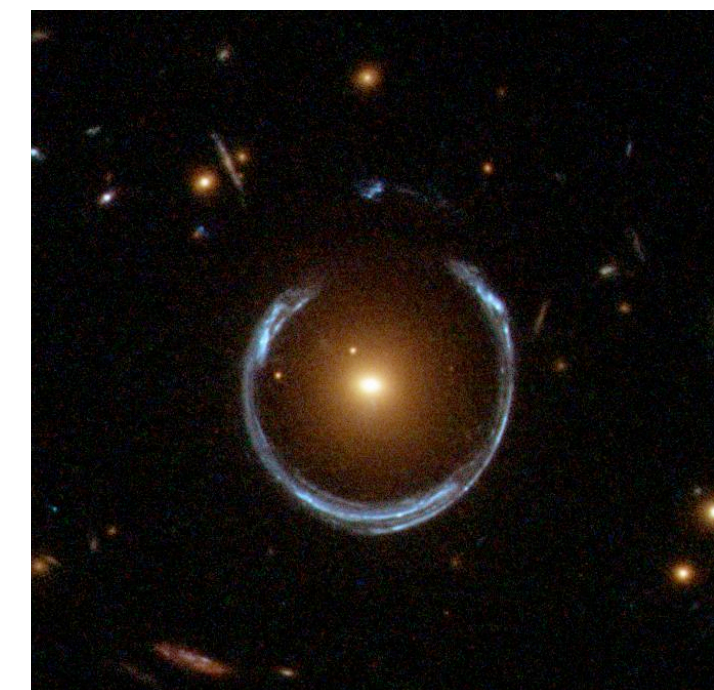
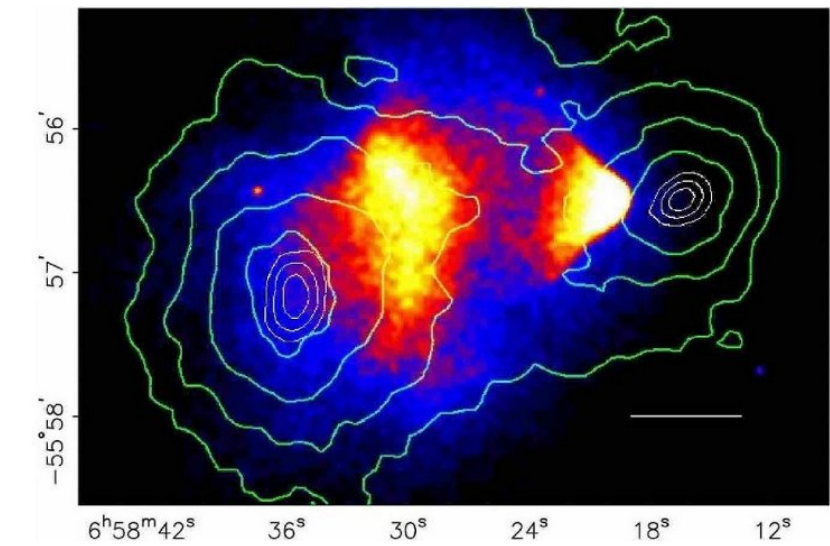
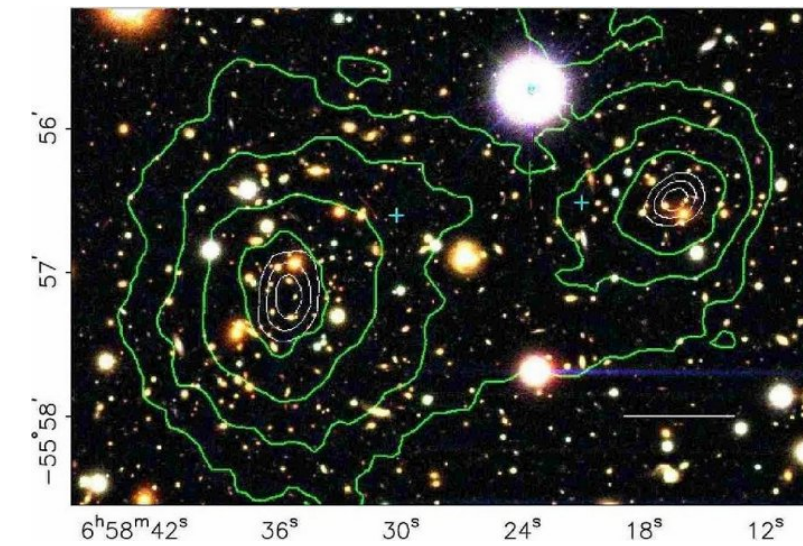
# Gravitational lensing

---

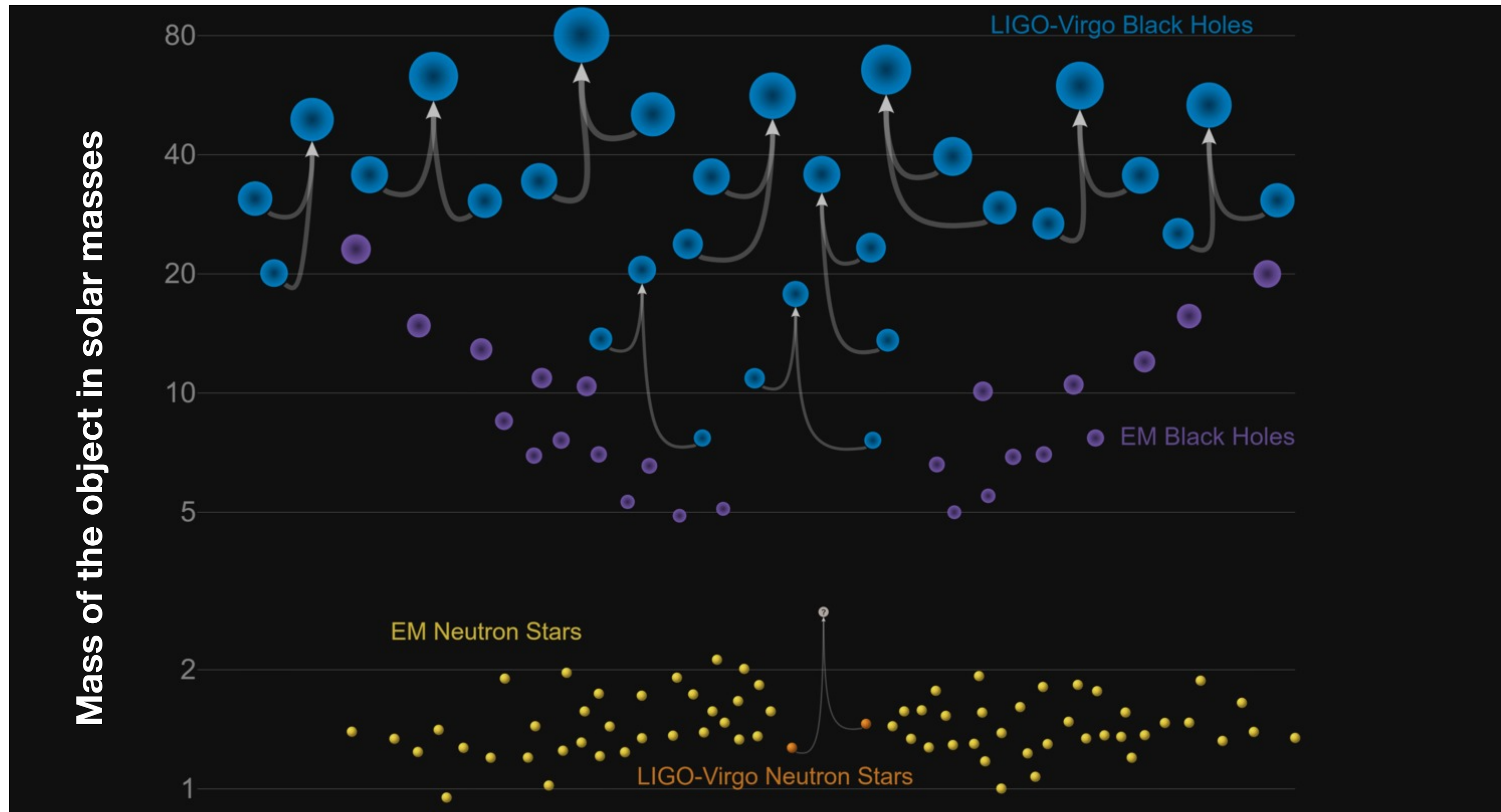


# Gravitational lensing: A powerful tool for astronomy

- Evidence of dark matter.
- Mapping out dark matter distributions.
- Constraining the fraction of compact dark matter (MACHOs).
- Detection of exoplanets.
- Estimation of cosmological parameters.
- ....



# 2015: Birth of gravitational-wave astronomy



[LIGO-Virgo/Frank Elavsky/Northwestern]

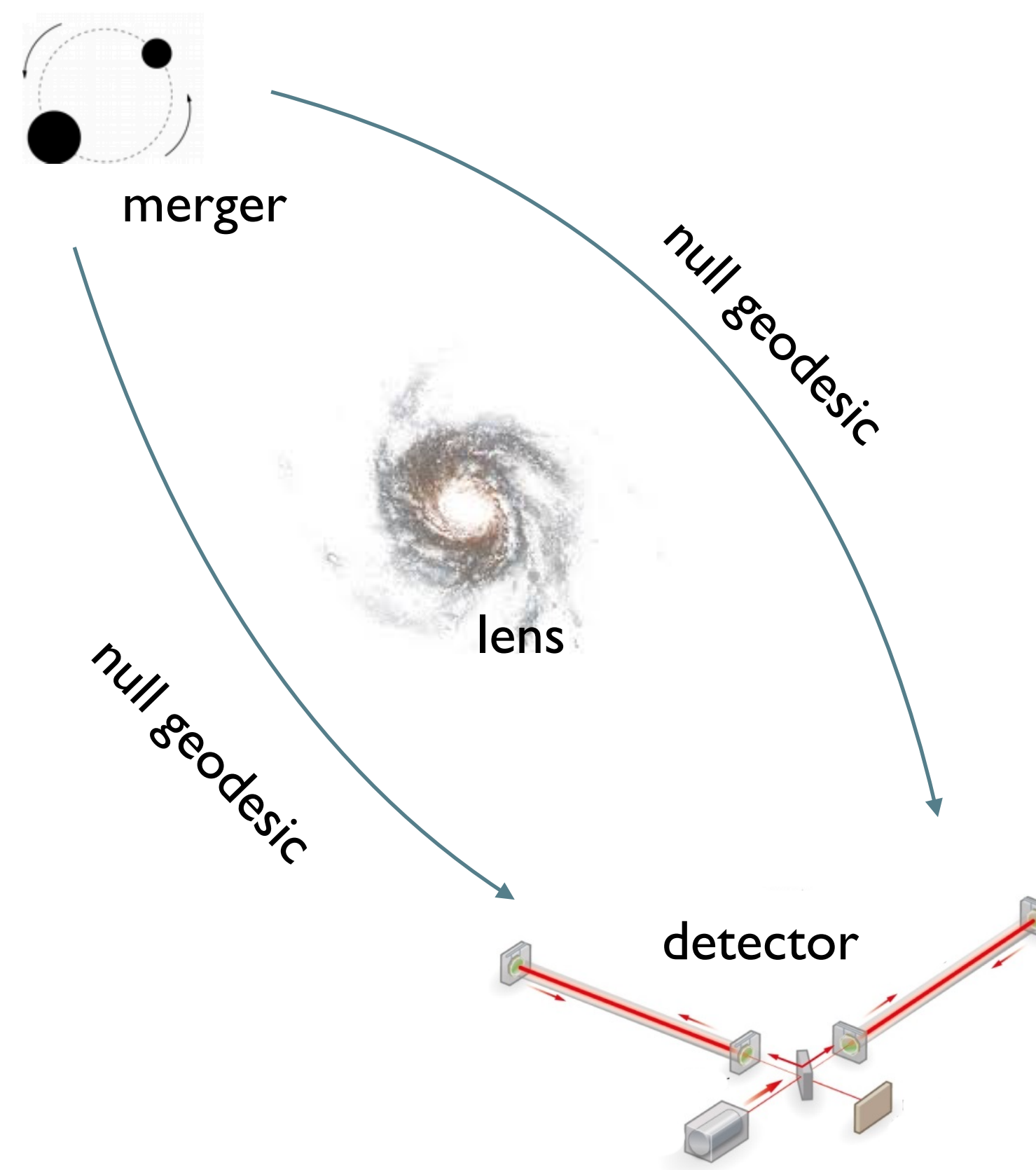
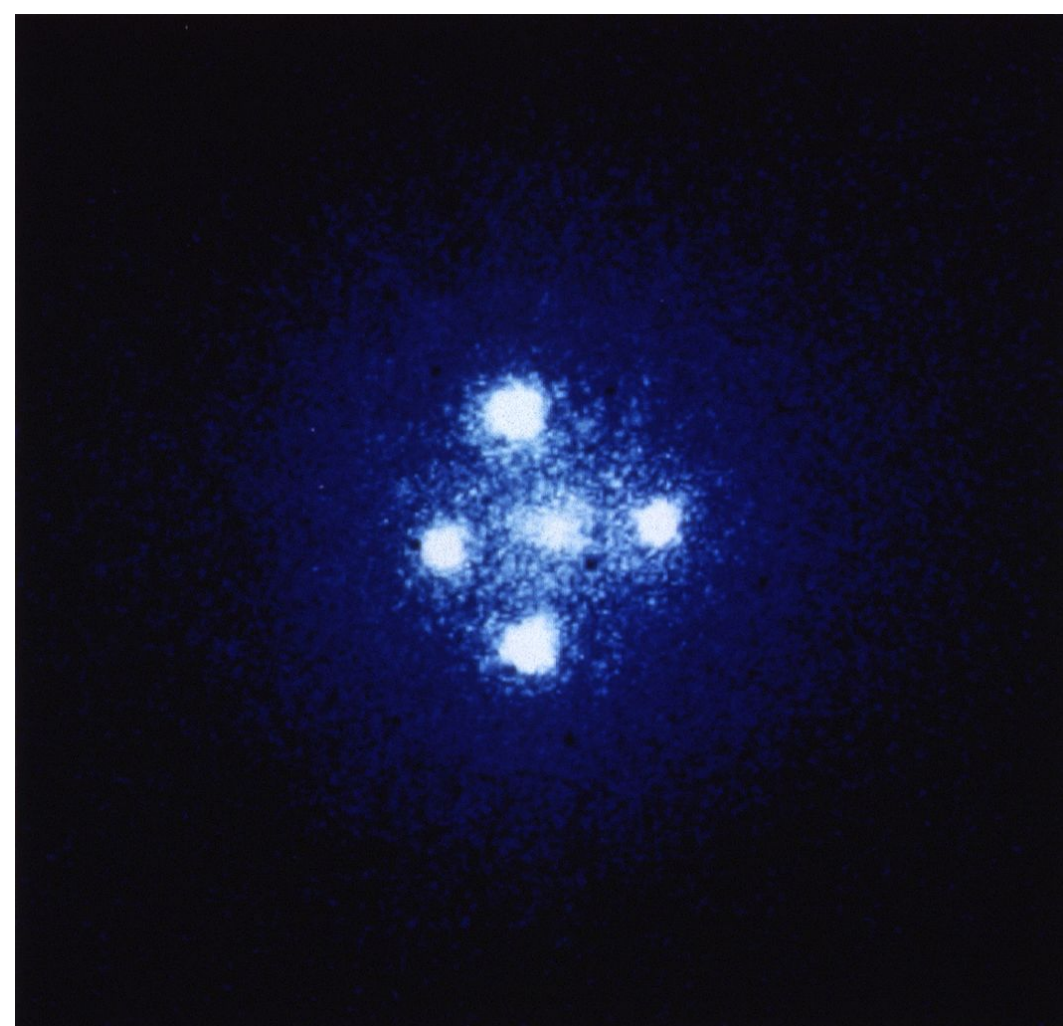
10 BBH + 1 BNS detections from the first two observing runs (LVC analysis)

Additional events from independent analyses of the data.

~50 BBH, BNS and NSBH candidates from the ongoing third observing run (1 BNS published) <sup>6</sup>

# Gravitational lensing of GWs: A new frontier

- Small fraction ( $\sim 0.4\%$ ) of detectable BBH mergers could be strongly lensed by intervening galaxies  $\Rightarrow$  multiple images, separated by hours to weeks.



# Gravitational lensing of GWs: A new frontier

---

- Small fraction ( $\sim 0.4\%$ ) of detectable BBH mergers could be strongly lensed by intervening galaxies  $\Rightarrow$  multiple images, separated by hours to weeks.

**A hundred years after Eddington's observation of the bending of light, we are in the verge of detecting the gravitational bending of GWs!**



# Identifying lensed images is challenging

---

- Need to identify pairs of lensed events among thousands of unrelated events.

# Identifying lensed images is challenging

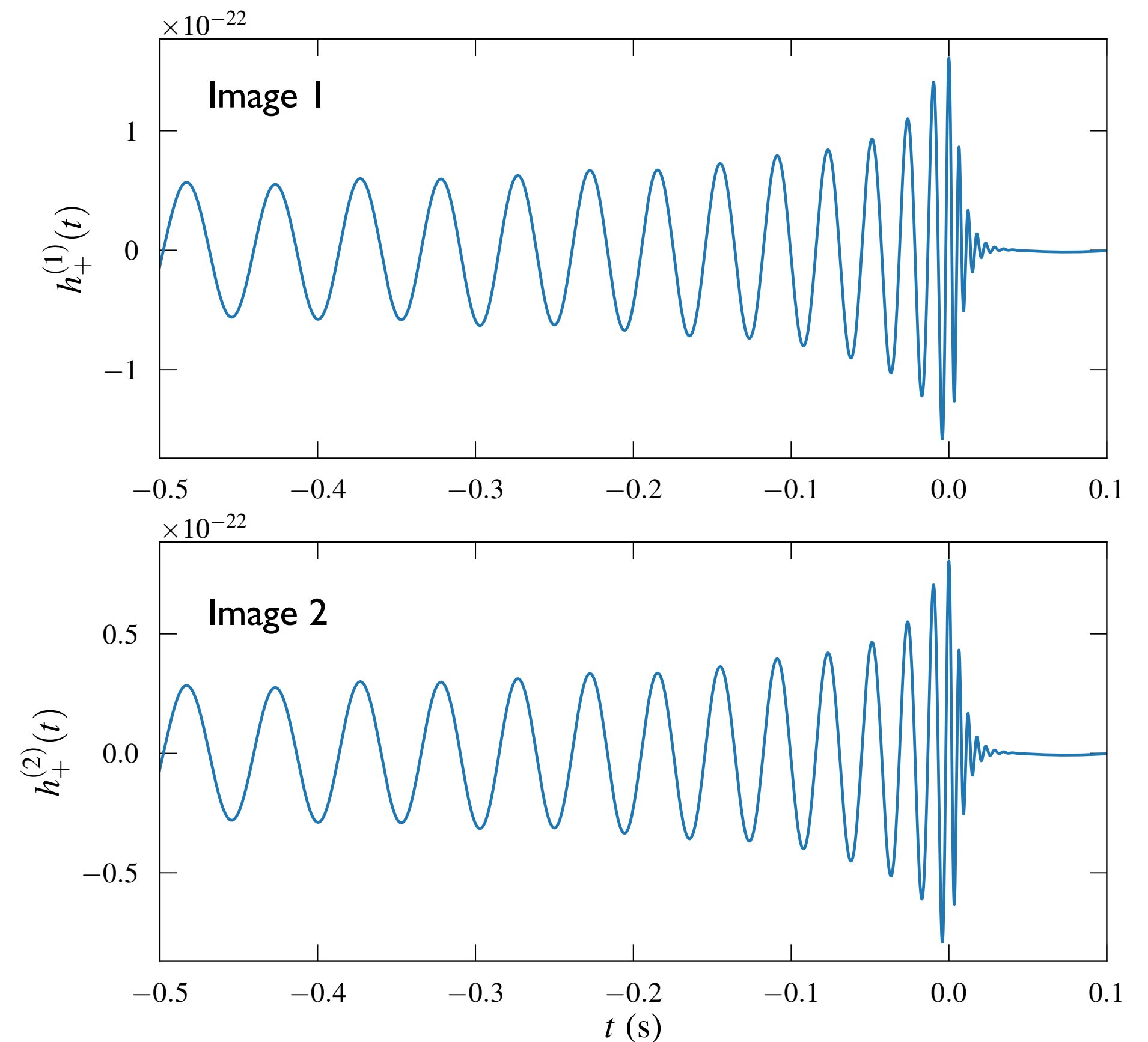
- Need to identify pairs of lensed events among thousands of unrelated events.
- Galaxies are the dominant lenses. Geometric optics regime ( $\lambda_{\text{GW}} \ll M_{\text{lens}}$ ). Signals are only magnified; without affecting their shape.

$$h_{+, \times}^i(t) = \sqrt{\mu^i} h_{+, \times}(t)$$

Lensed polarizations  
from multiple images

Lensing  
magnification

Original polarizations



# Identifying lensed images is challenging

---

- Need to identify pairs of lensed events among thousands of unrelated events.
- Galaxies are the dominant lenses. Geometric optics regime ( $\lambda_{\text{GW}} \ll M_{\text{lens}}$ ). Signals are only magnified; without affecting their shape.

$$h_{+, \times}^i(t) = \sqrt{\mu^i} h_{+, \times}(t)$$

Distance is degenerate with magnification. Estimated distance will be biased; so will be the estimated redshift and source-frame masses.

$$d_{\text{obs}} = \frac{d_L}{\sqrt{\mu}}$$

$$z_{\text{obs}} = z(d_{\text{obs}})$$

$$M_{\text{obs}} = \frac{M^z}{1 + z_{\text{obs}}}$$

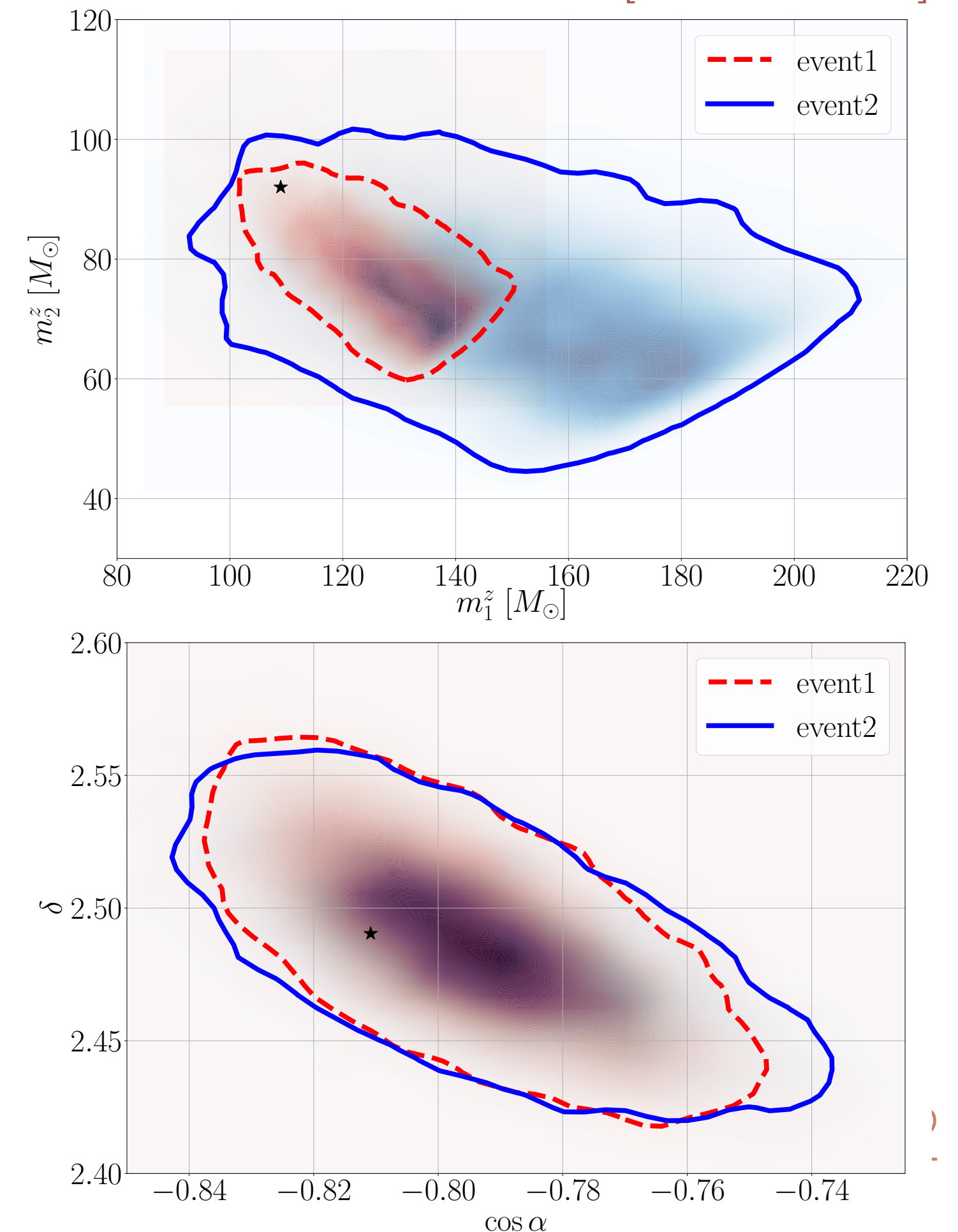
# Identifying lensed images is challenging

- Bayesian model selection: From a pair of events compute the odds ratio of the two hypotheses (lensed vs unlensed)

$$O_{\text{U}}^{\text{L}} := \underbrace{\frac{P(\mathcal{H}_{\text{L}})}{P(\mathcal{H}_{\text{U}})}}_{\text{Prior odds}} \underbrace{\int d\vec{\theta} \frac{P(\vec{\theta}|d_1) P(\vec{\theta}|d_2)}{P(\vec{\theta})}}_{\text{Lensing Bayes factor}}$$

Posterior distributions of binary parameters (except the distance) estimated from lensed images will be consistent

[Haris et al 2018]

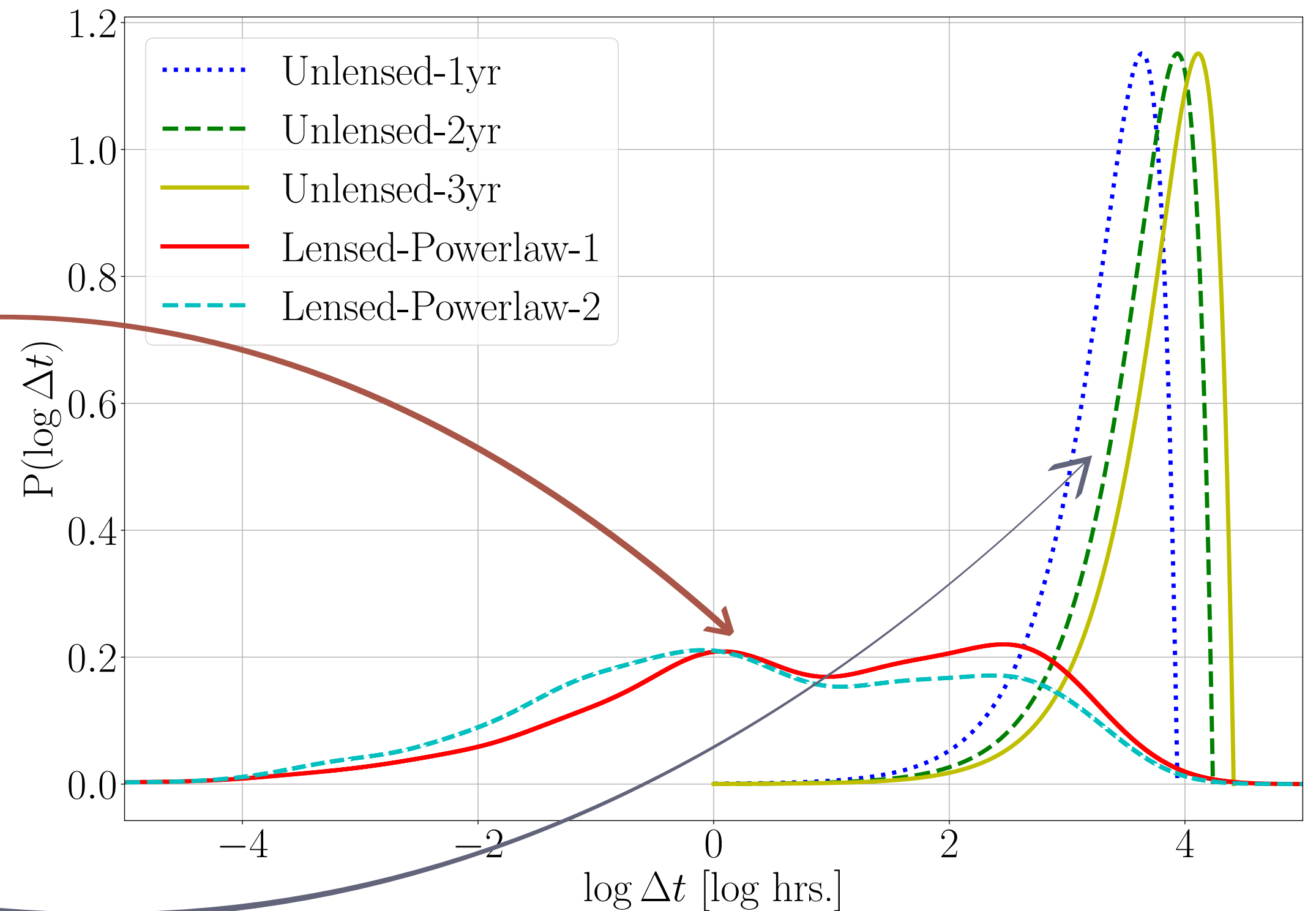


# Identifying lensed images is challenging

[Haris et al 2018]

- Bayesian model selection:  
Additional information from the  
time delay between events.

$$\mathcal{R}_U^L = \frac{P(\Delta t_0 | \mathcal{H}_L)}{P(\Delta t_0 | \mathcal{H}_U)}$$

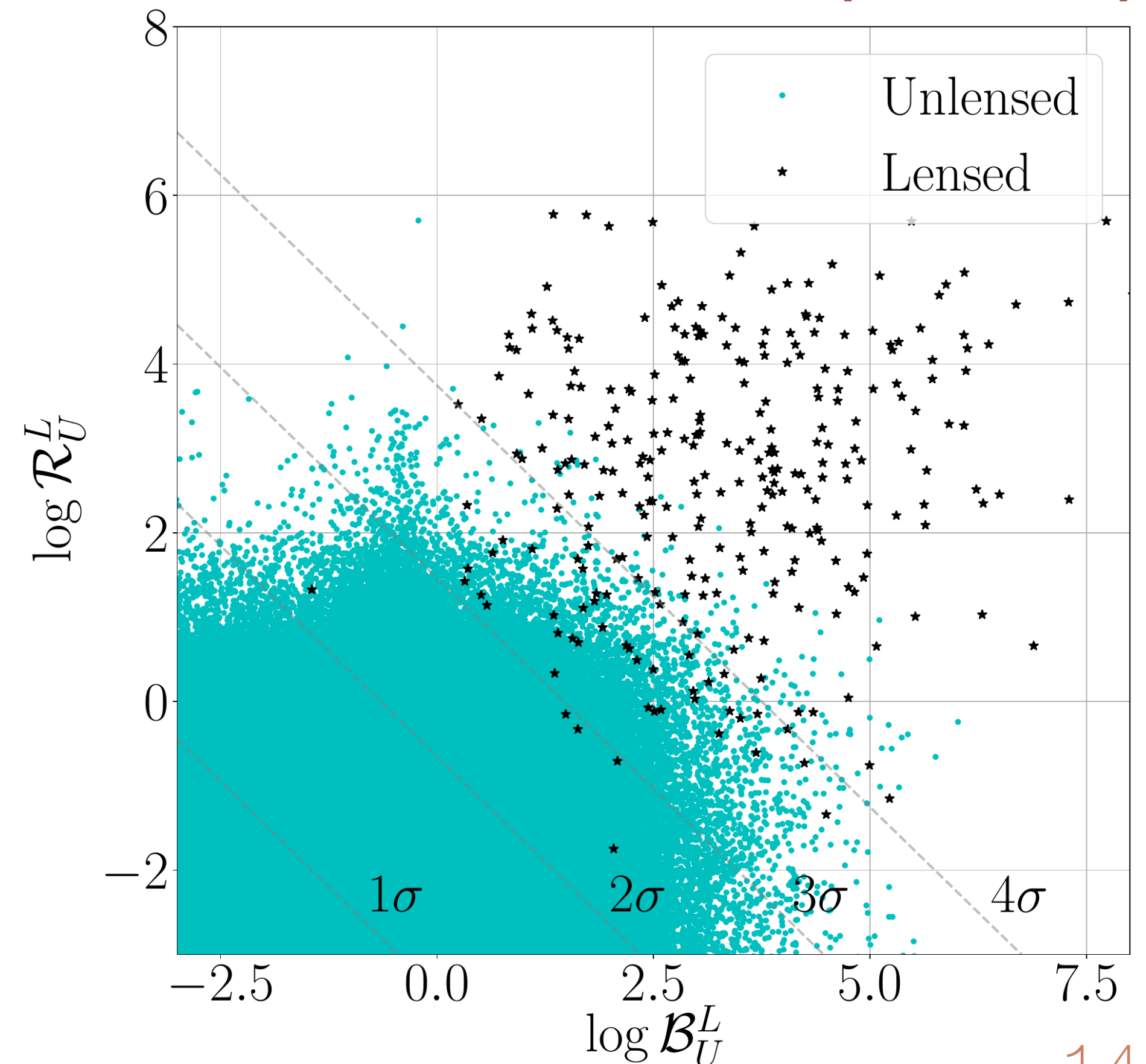


Prior distributions of time delay expected  
from lensed and unlensed events.

# Identifying lensed images is challenging

- Combine the Bayes factors to discriminate between lensed & unlensed events.
  - ~66% of the lensed events can be identified with a false alarm probability of  $10^{-5}$ .

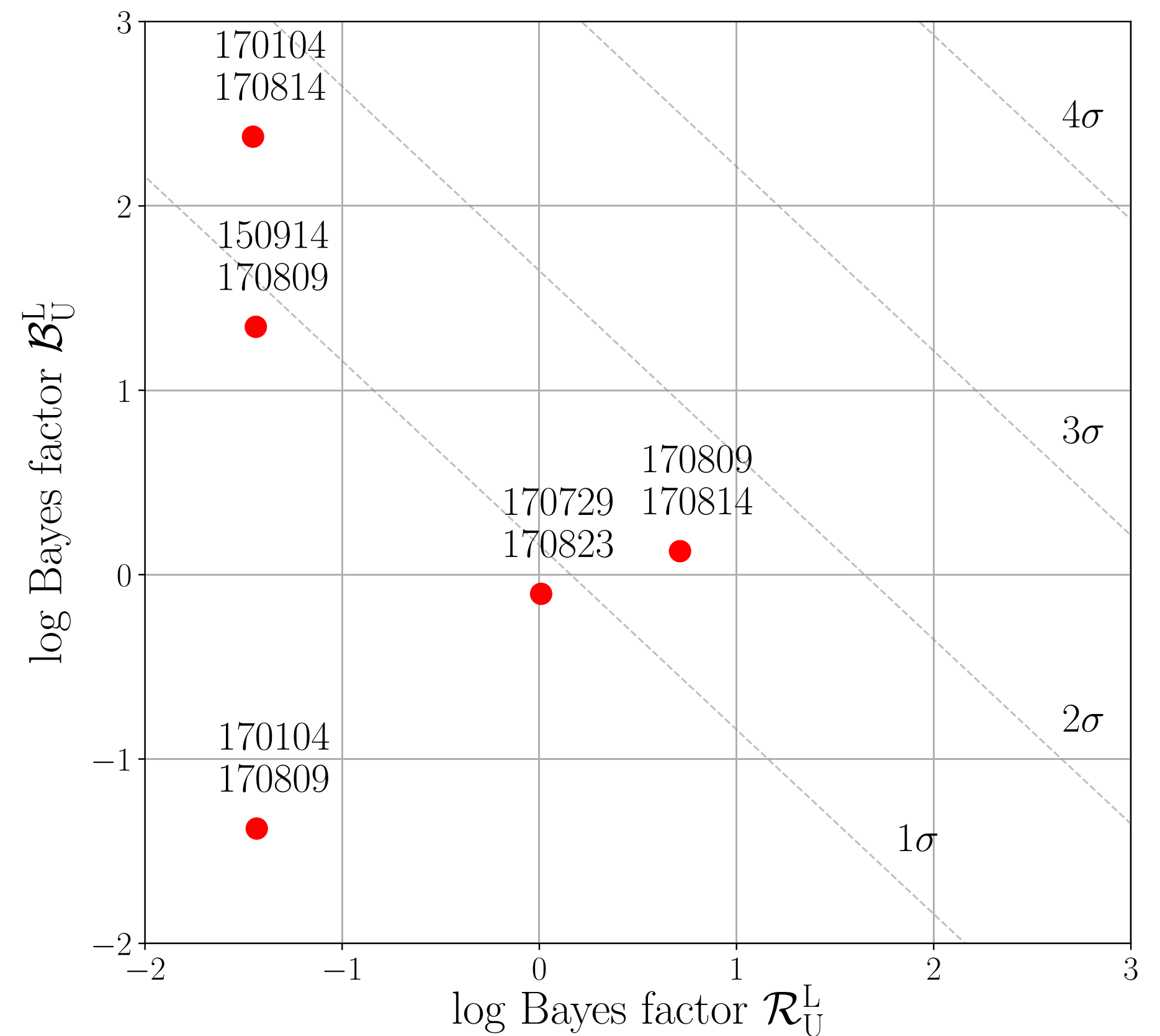
[Haris et al 2018]



# No evidence of lensing... yet.

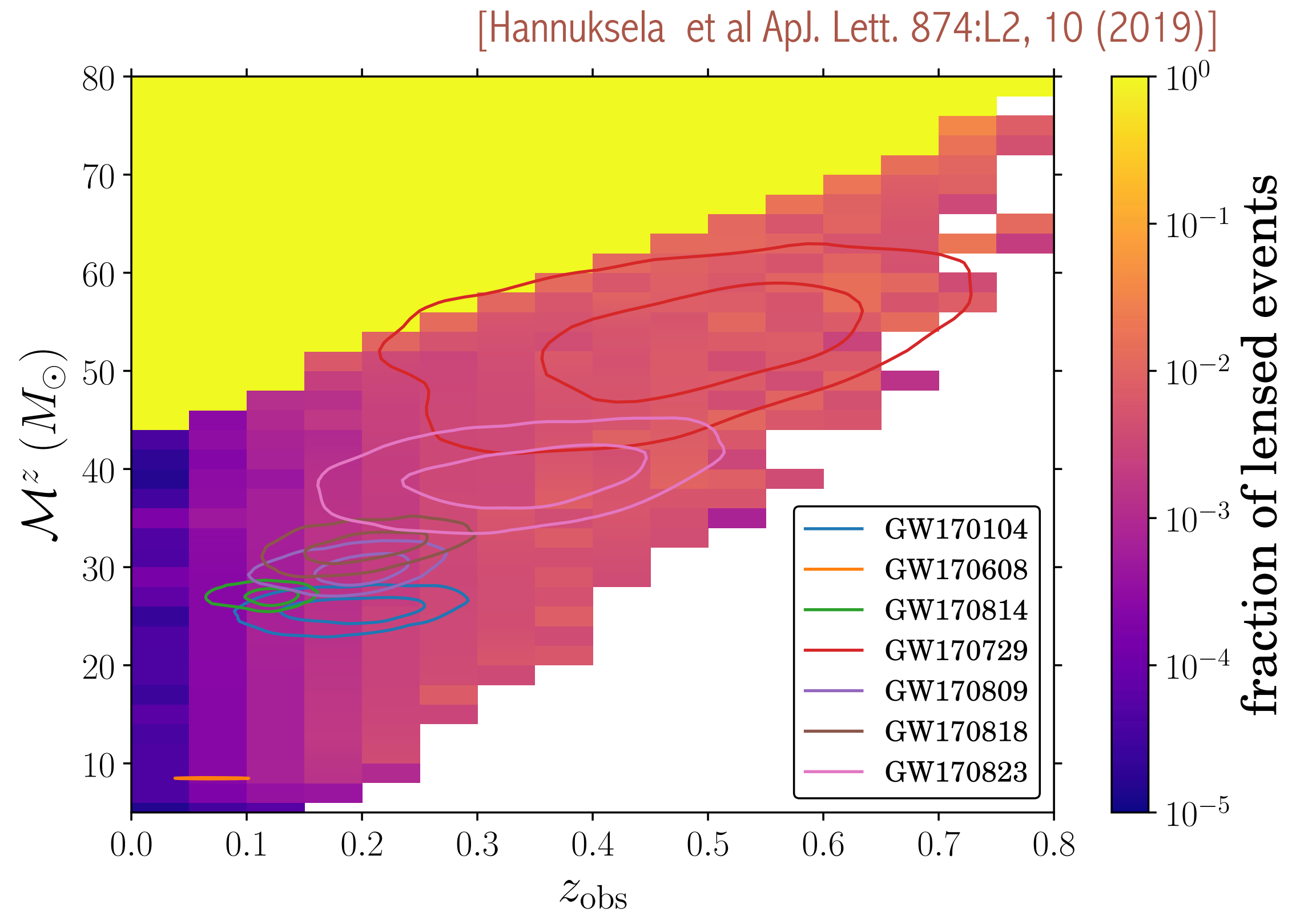
- Performed the first search for lensing effects in the binary BH observations during the first 2 observing runs.
  - No evidence of multiple images.

[Hannuksela et al ApJ. Lett. 874:L2, 10 (2019)]



# No evidence of lensing... yet.

- Performed the first search for lensing effects in the binary BH observations during the first 2 observing runs.
- Observed source properties are consistent with expectations without invoking lensing magnification.





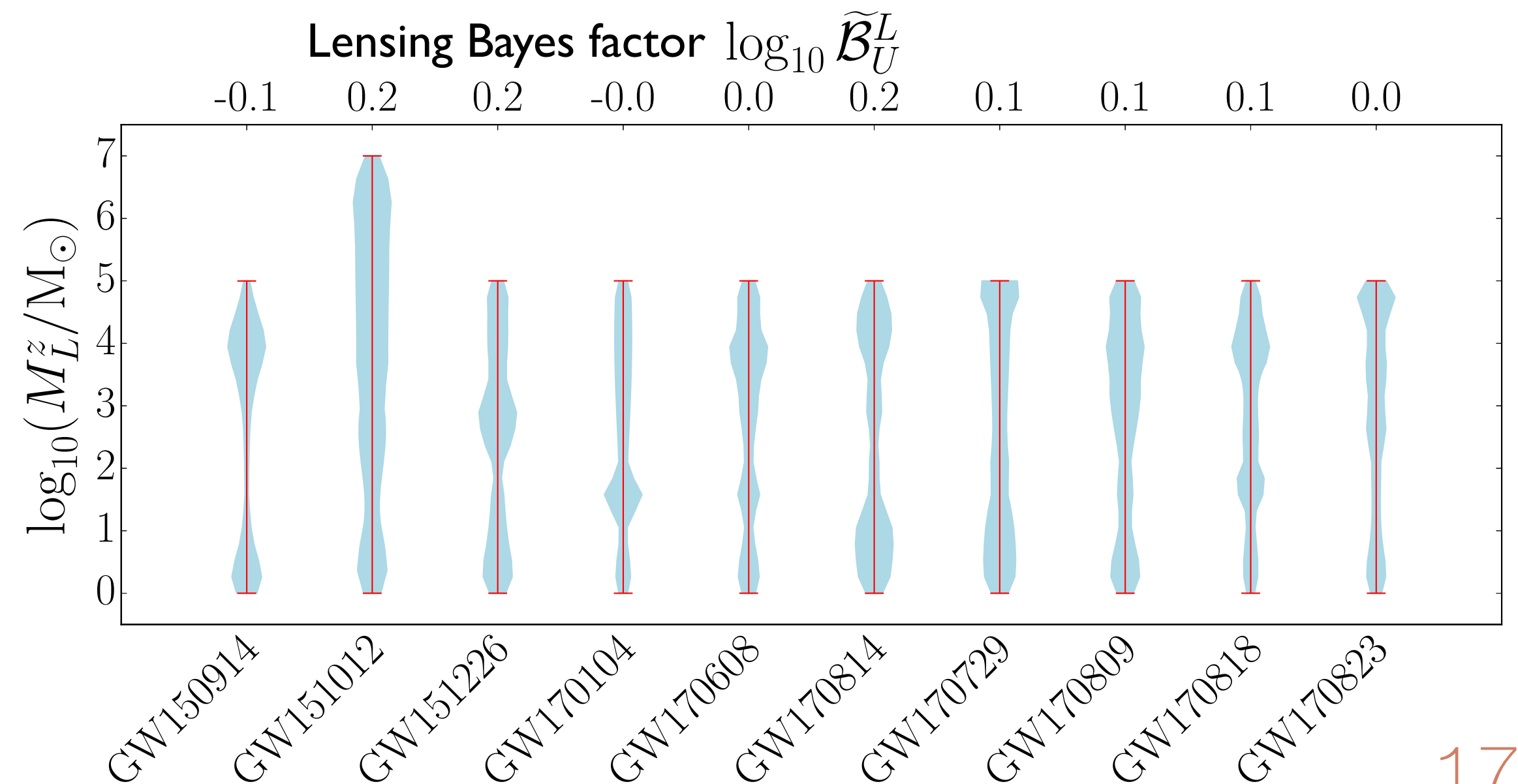
# No evidence of lensing... yet.

- Performed the first search for lensing effects in the binary BH observations during the first 2 observing runs.

[Hannuksela et al ApJ. Lett. 874:L2, 10 (2019)]

When  $\lambda_{\text{GW}} \sim M_{\text{lens}}$  lensing can produce wave optics effects. Expected deformation in the GW signals can be modeled for simple lens models (e.g point mass).

- No evidence of wave optics effects due to lensing in the observed signals.



No evidence of lensing... yet.

- Performed the first search for lensing effects in the binary BH observations during the first 2 observing runs

SCIENTIFIC  
AMERICAN



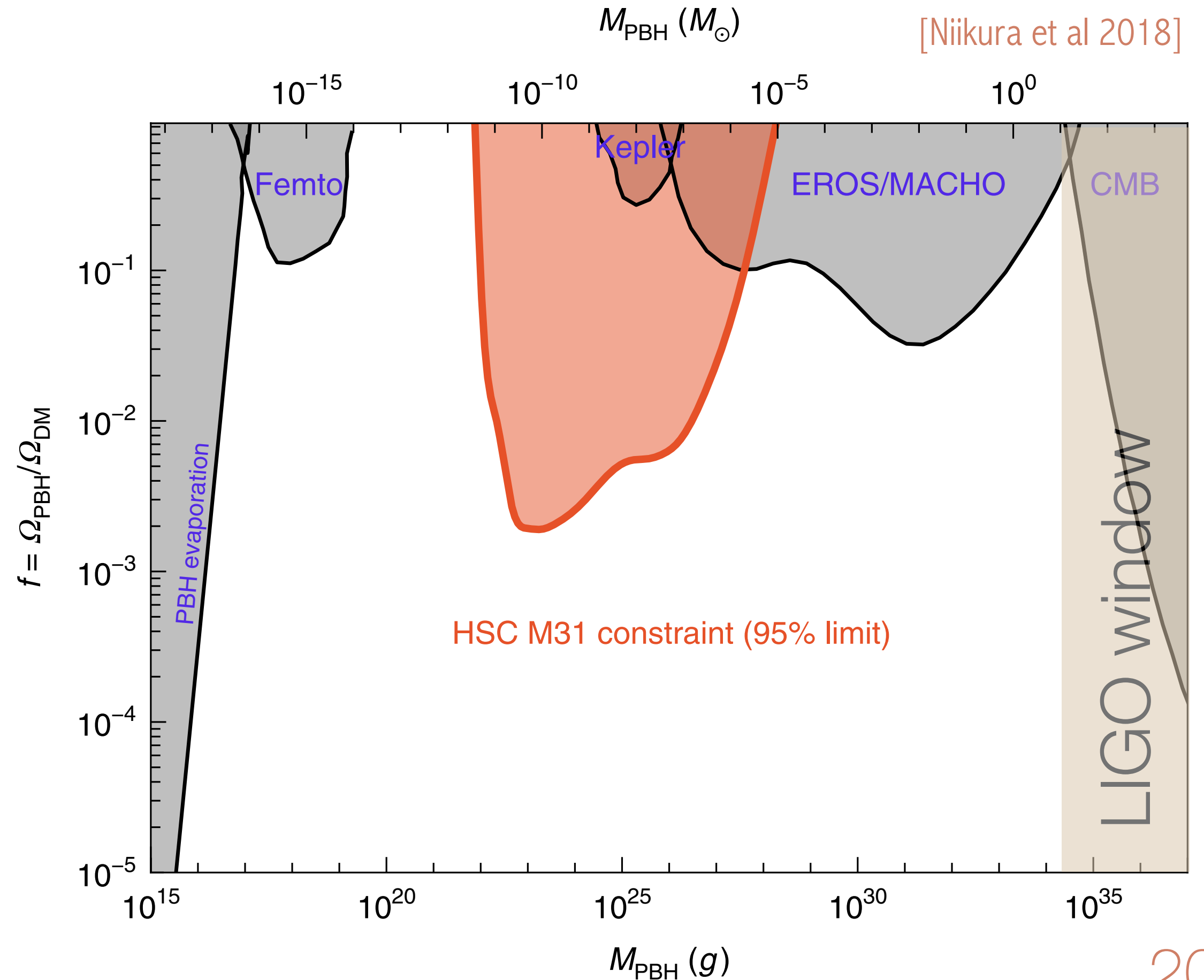
# Gravitational lensing of GWs: A new frontier

---

- LIGO-Virgo expected to detect 100s of mergers in the next few years. First detection of lensed GWs around the corner.
- Precise (sub-galaxy) localization of mergers from lensed images from the observed time delay and magnification ratio [Mehta et al, In prep]
- Accurate extraction of the polarizations. Are polarizations consistent with GR predictions? [Goyal et al, In prep]

# Gravitational lensing of GWs: A new frontier

- If the lens mass is comparable to the GW wavelength ( $10\text{-}10^5 M_\odot$ )  $\Rightarrow$  wave optics signatures in the observed signal.
- Constrain the fraction of dark matter in the form of compact objects (even from null observations) [Ganguly et al, In prep]



# Summary

---

- Gravitational lensing observations begun as a means of testing the validity of GR.
- Eventually it became a powerful probe of astrophysics and cosmology.
- Similar story of GWs. Now becoming a unique branch of astronomy.
- More fascinating observations are expected in the near future, e.g., the gravitational lensing of GWs — doubly Einstein!