

# Quantum Gravity & Quantum Information

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# Quantum Gravity & Quantum Information

[1] Two Puzzles in Black Hole Information

QG



[2] Page Curve and Recent Results

QI



[3] Typical Black Hole Entanglement Entropy

QG

# 1 Two Puzzles in Black Hole Information

i) What is the origin of black hole entropy?

$$S_{\text{BH}} = \frac{A_{\text{BH}}}{4G\hbar}$$

ii) What is the fate of black hole information?

$$A_{\text{BH}}(t) \rightarrow 0 \quad \text{evaporation}$$

$$\left\{ \begin{array}{l} S_{\text{BH}}(t) \rightarrow 0 \\ S_{\text{rad}}(t) \rightarrow \infty \end{array} \right.$$

# (i) Black Hole Entropy

(Bekenstein, Hawking '74)

## • Thermodynamic derivation

Assumptions:

- Equilibrium
- Vacuum & Isolated
- here non-rotating for simplicity

a) GR Horizon Area  $A_{BH} = 4\pi R^2$   
with  $R = 2GM/c^2$

b) QFT Temperature  $k_B T = \frac{\hbar c}{4\pi R}$

c) Thermodyn. Entropy  $\delta S = \frac{\delta Q}{T}$



Entropy

$$\delta S_{BH} \stackrel{(c)}{=} \frac{\delta(Mc^2)}{T} \stackrel{(a), (b)}{=} k_B \frac{\delta A_{BH}}{4G\hbar/c^3}$$

\* Puzzle:

Pure State QM

• For Non-Grav Systems: Thermodynamics ← Stat. Mech ← Entropy from Entanglement

• For BH:

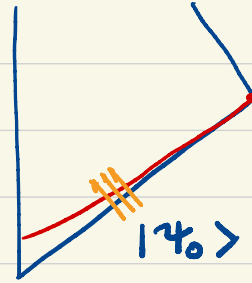
$S_{BH}$  ← ? ← ?

## ii) Fate of Black Hole Information

(Hawking '76)  
(Page '93)

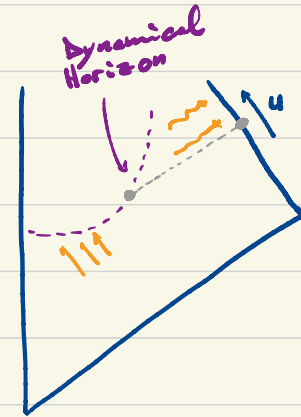
- Initial Conditions:

pure state of grav. & matter



- Isolated, unitary evolution

BH formation  $S_{BH}$



- Evaporation

$S_{rad}(u)$  increases  
 $S_{BH}(u)$  decreases

$S_{tot}(u) = S_{BH}(u) + S_{rad}(u)$  increases (GSL)

- Hawking Information Puzzle:  $M_B(u) \rightarrow 0$ ,  $|\psi_0\rangle \rightarrow \rho_{rad}$  ?

- Page Information Puzzle:  $M_B(u) < \frac{M_B(u_0)}{\sqrt{2}}$ , profile of  $S_{rad}(u)$  ?

# Page Curve: Qubit Model of Evaporation

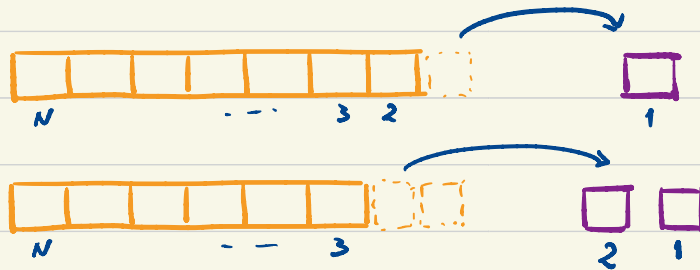
(Page 193)

- $N$  qubits in a pure state

$$|\psi\rangle = \sum_{i_1 \dots i_N = \pm 1} c_{i_1 \dots i_N} |i_1\rangle \dots |i_N\rangle$$

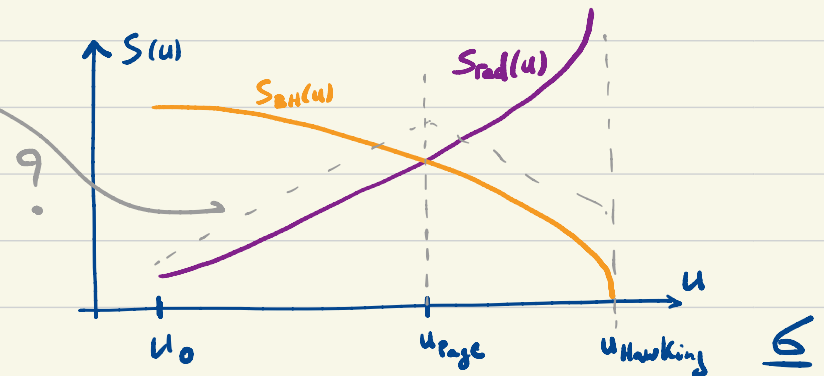
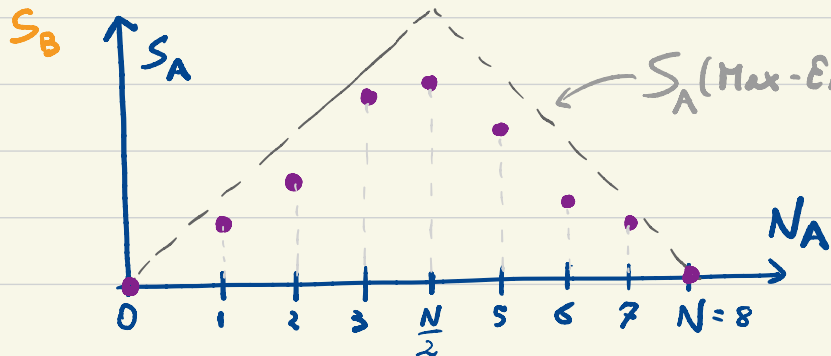


- Extract one qubit at a time



- Entanglement Entropy

$$S_A(|\psi\rangle) = -\text{Tr}_A(\rho_A \log \rho_A) \quad \text{with} \quad \rho_A = \text{Tr}_B |\psi\rangle\langle\psi|$$



BH formation



BH evaporation

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[3] Typical Black Hole Entanglement Entropy

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(i) Typical Entropy of Random States

(ii) Time Evolution and Thermalization

(iii) Typical Entropy with Constraints

(iv) Typical Entropy of Sub-Algebras

## [2] (i) Page Curve and Typical Entropy of Random States

• Hilbert Space  $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$

$$d_A = \dim \mathcal{H}_A \\ d_B = \dim \mathcal{H}_B$$

• Random Pure State  $|\psi\rangle = U |\psi_0\rangle$

Random Unitary

from uniform probability distribution  $d\mu(U)$  (Haar Measure)

Reference State

• Entanglement Entropy  $S_A(|\psi\rangle)$

\* What is the probability of finding  $S_A$ ?

→ compute  $P(S_A) dS_A$

• Page '93: Average Entropy of a Subsystem  $\langle S_A \rangle = \int S_A(U|\psi_0\rangle) d\mu(U)$

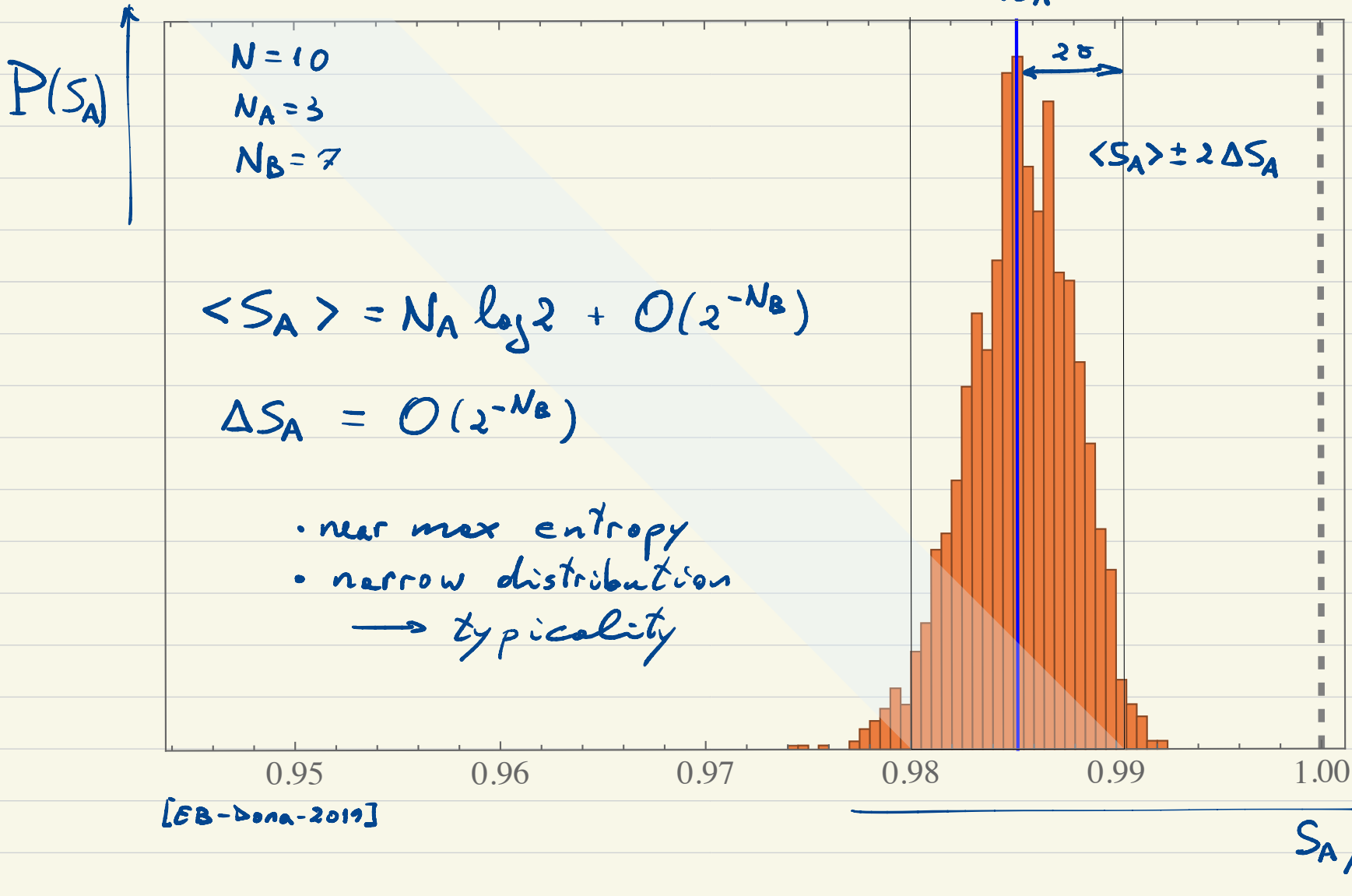
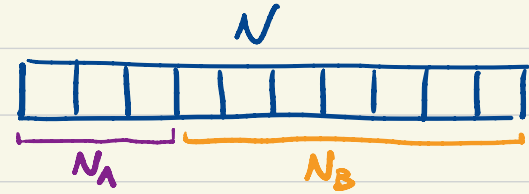
• Bianchi-Donà '19: Typical entropy  $\langle S_A \rangle \pm \Delta S_A$

from moments  $\mu_n = \int (S_A)^n P(S_A) dS_A$

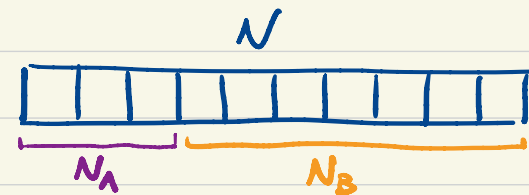


Example:  $N$  qubits

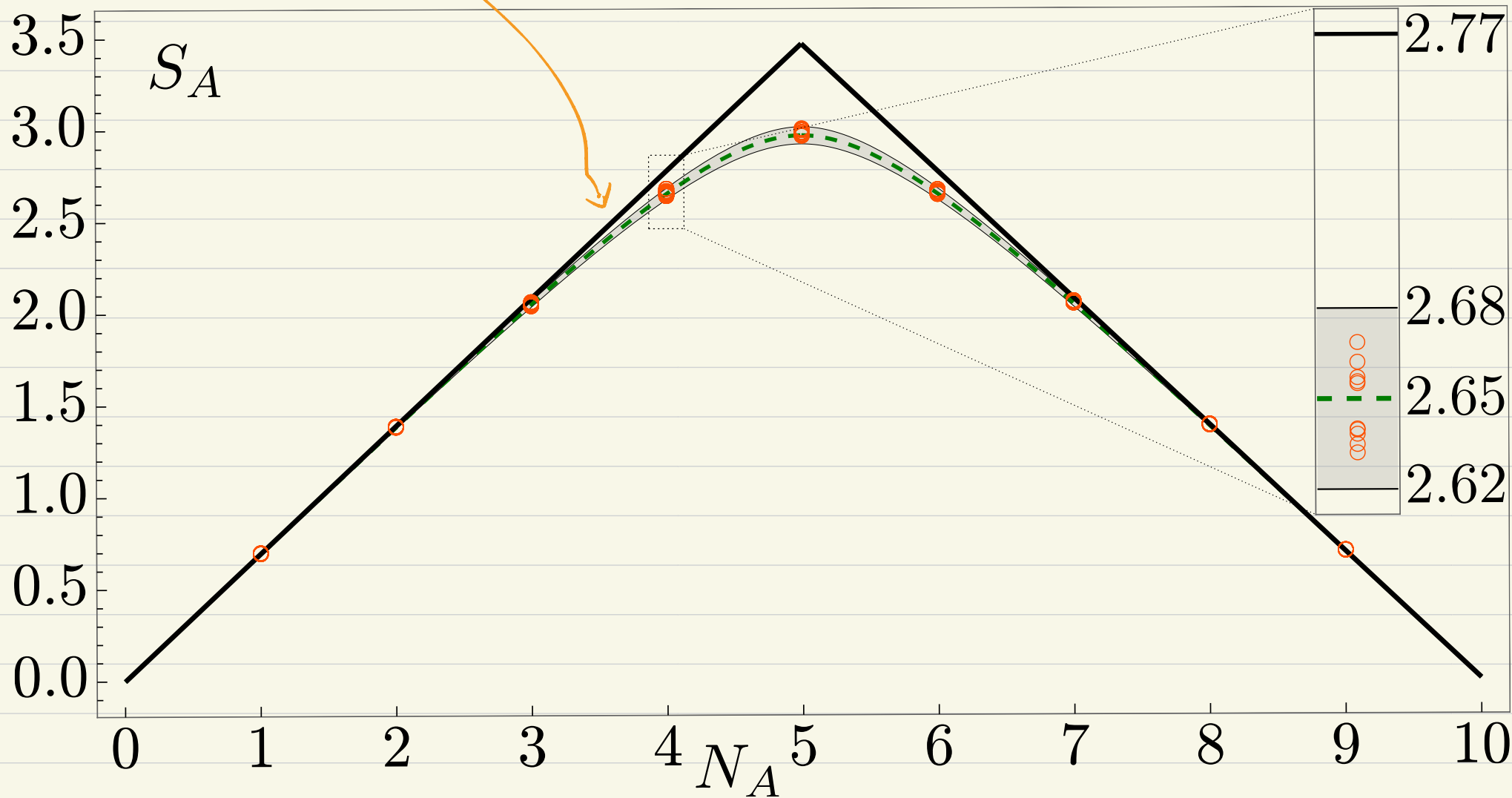
$$d_A = 2^{N_A}, \quad d_B = 2^{N_B}$$



Example:  $N$  qubits



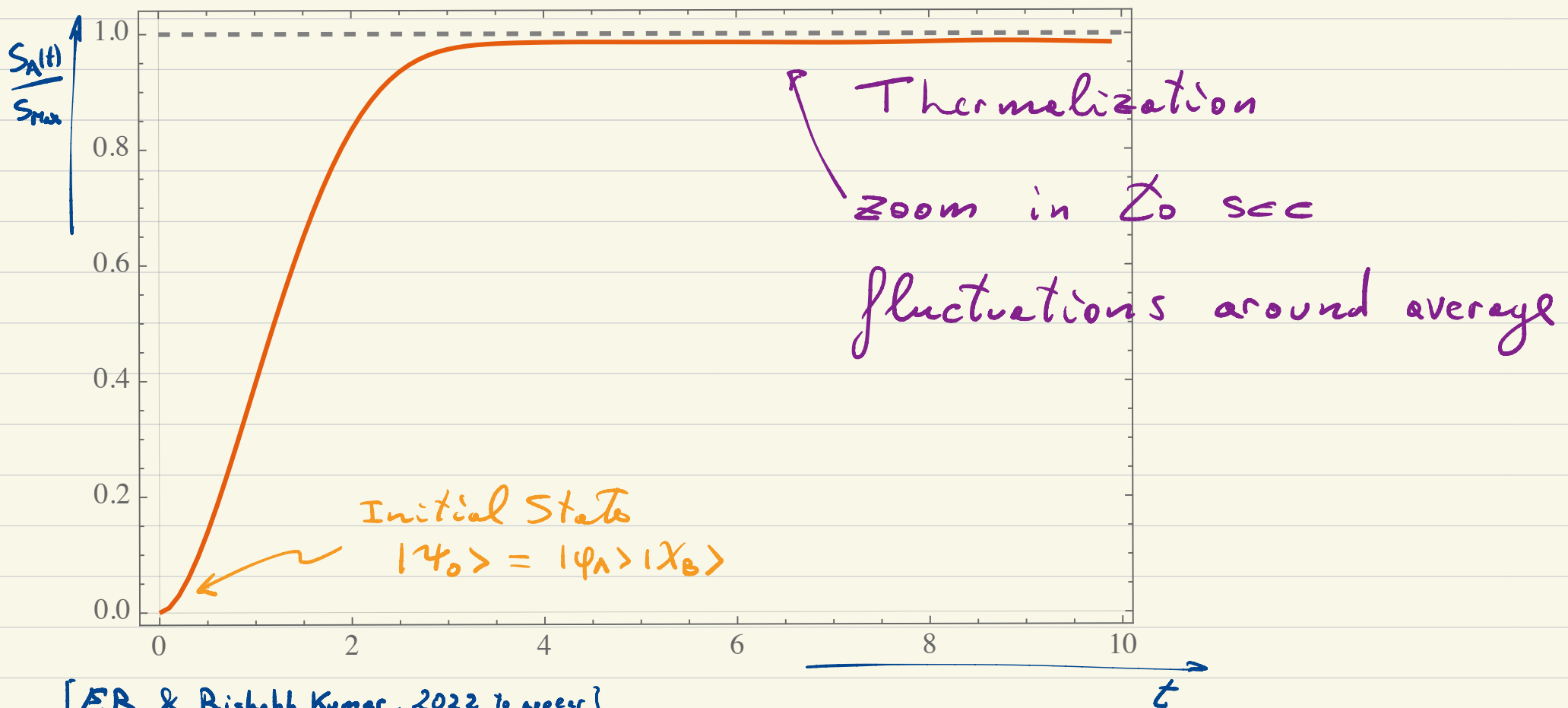
Page Curve



[EB-Done-2019]

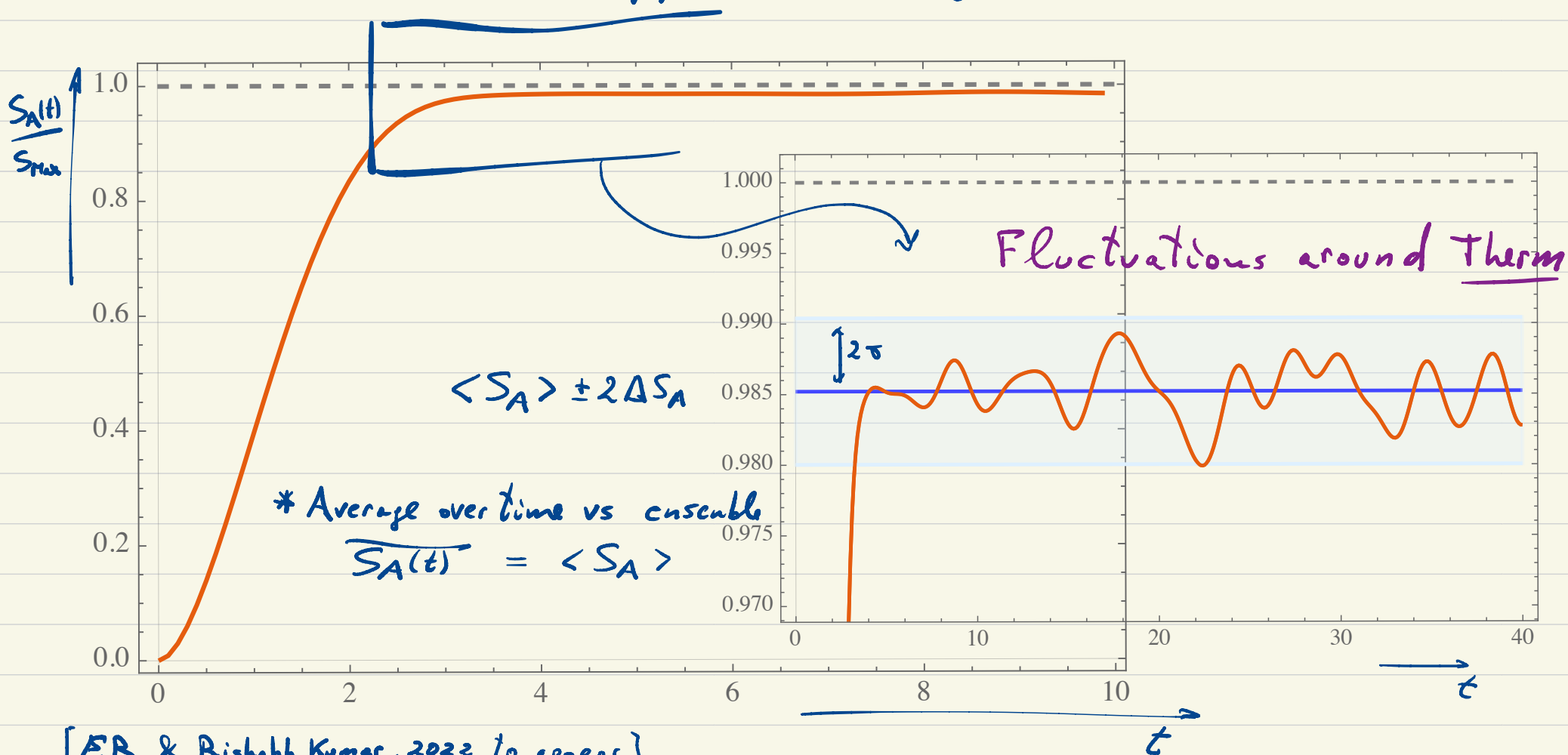
## ■ (ii) Time Evolution of the Entanglement Entropy

- Initial Factorized State  $|\psi_0\rangle = |\psi_A\rangle |\chi_B\rangle$
- Evolution with Random Hamiltonian  $H = \sum_{ij} H_{ij} |i\rangle\langle j|$
- Entanglement Entropy  $S_A(t)$  of  $|\psi_t\rangle = e^{iHt} |\psi_0\rangle$



# Time Evolution of the Entanglement Entropy

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- Evolution with Random Hamiltonian  $H = \sum_{ij} H_{ij} |i\rangle\langle j|$
- Entanglement Entropy  $S_A(t)$  of  $|\psi_t\rangle = e^{iHt} |\psi_0\rangle$



### ■ (iii) Typical Entanglement Entropy of Random States with Constraints

• Hilbert Space  $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$

• Constraint  $|\psi_E\rangle \in \mathcal{H}_E \subset \mathcal{H}$

e.g. • Hamiltonian  $H = H_A + H_B$

• Energy Eigen-space  $H|\psi_E\rangle = E|\psi\rangle$

• Decomposition as direct sum

$$\mathcal{H}_E = \bigoplus_E (\mathcal{H}_{A,E} \otimes \mathcal{H}_{B,E-E})$$

• Entanglement Entropy  $S_A(|\psi_E\rangle)$

• Probability distribution  $P_E(S_A) dS_A$

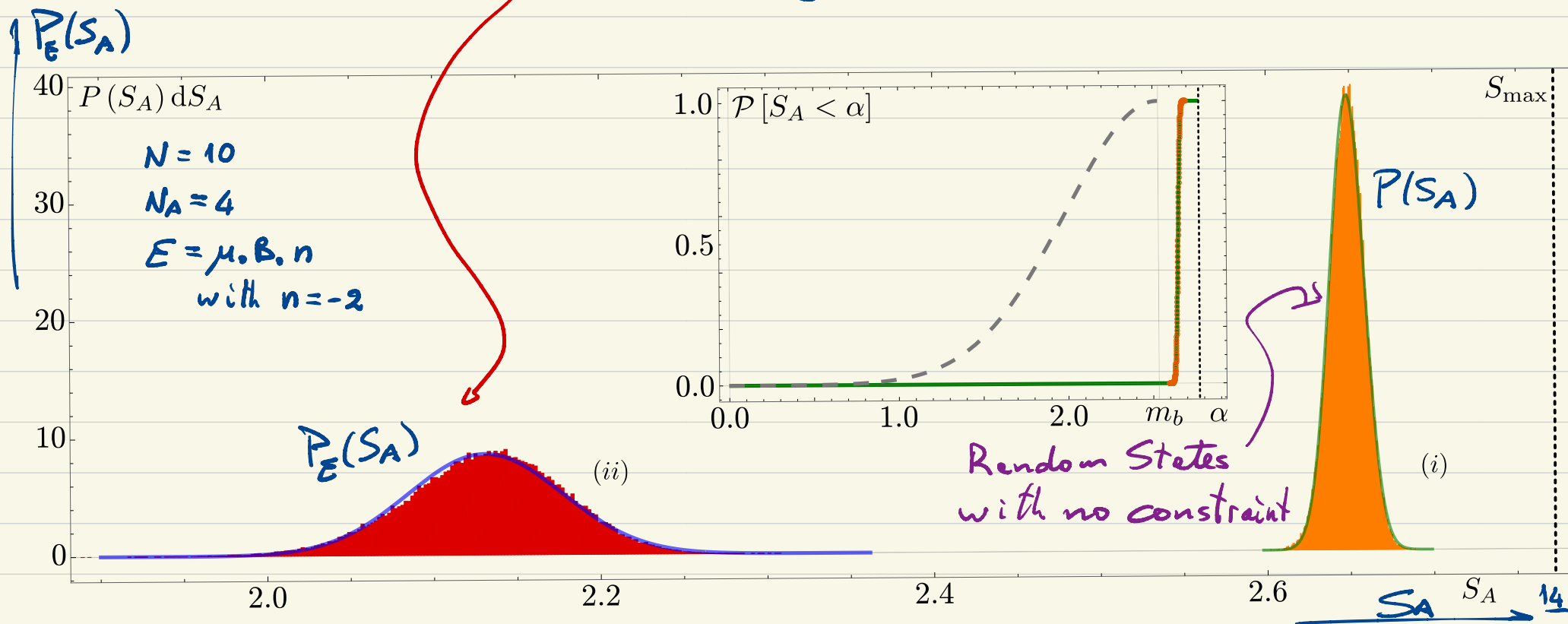
EB-Donà, 2019

EB-Hackl-Kieburg, 2021

EB-Hackl-Kieburg-Rigol-Vidmar, 2021

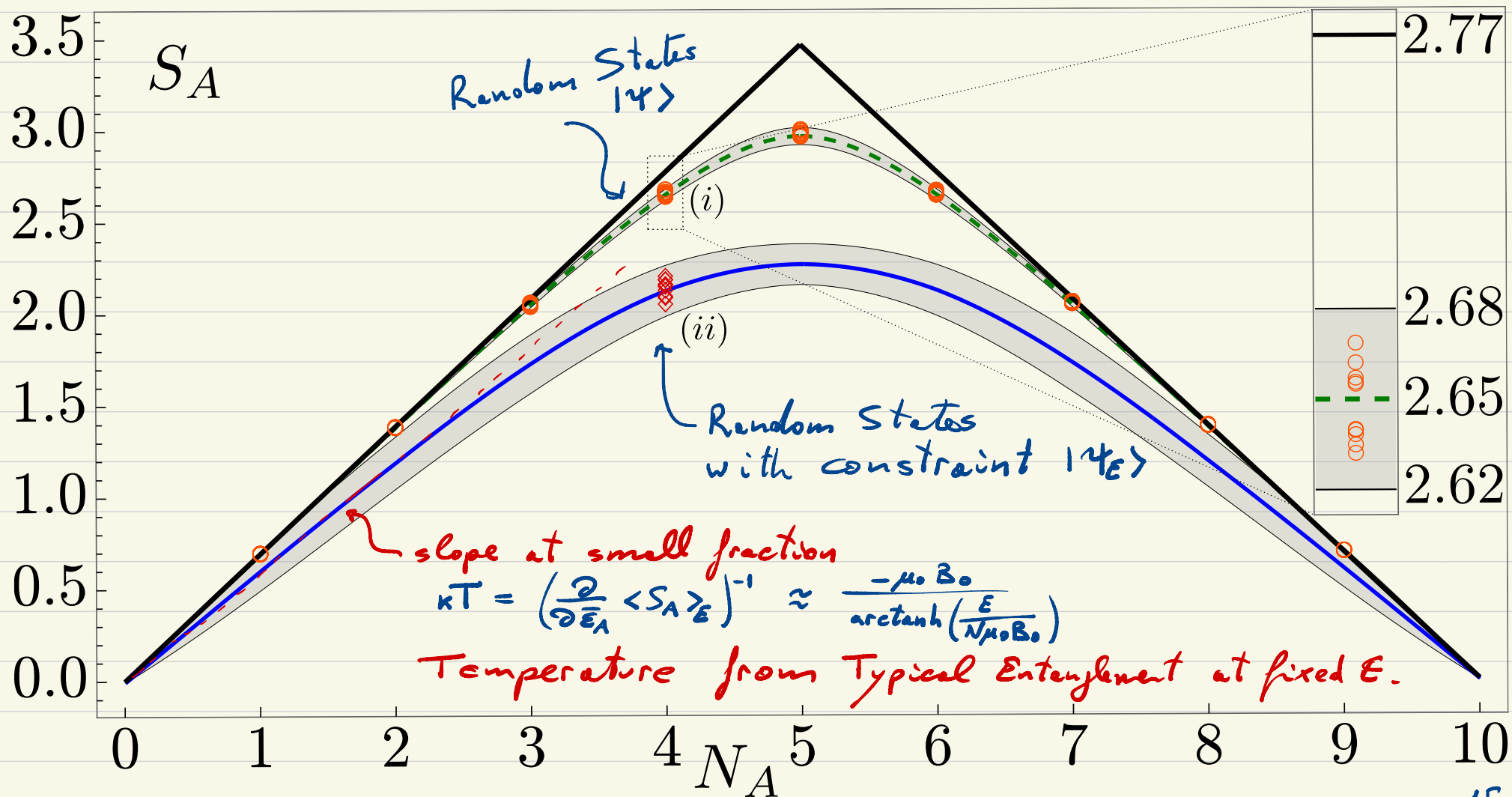
Example: typical entanglement in a paramagnet

- $N$  spins in a uniform magnetic field  $\vec{B}$
- Hamiltonian  $H = \sum_{i=1}^N \mu_0 \vec{B} \cdot \vec{S}_i$
- Random Eigenstate  $|\psi_E\rangle$  of energy  $E$
- Probability of finding entanglement entropy  $S_A$



Example: typical entanglement in a paramagnet

- Random Eigenstate  $|\psi_E\rangle$  of energy  $E$
- Page Curve



## ■ (iv) Entanglement Entropy of Sub-Algebras of Observables

• Generally,  $\mathcal{H} \neq \mathcal{H}_A \otimes \mathcal{H}_B$ , the Hilbert space does not come with a decomposition in subsystems

• Physically, measurements probe only part of the system

→ sub-algebra of observables  $\mathcal{O}_A \in A_A \subset \mathcal{A}$

→ not measured: commutant  $A_B = \{ \mathcal{O}_B \in \mathcal{A} \mid [\mathcal{O}_B, \mathcal{O}_A] = 0 \forall \mathcal{O}_A \}$

→ center  $\mathcal{Z} = A_A \cap A_B$

• Decomposition of the Hilbert Space

$$\mathcal{H} = \bigoplus_{\xi} \left( \mathcal{H}_A(\xi) \otimes \mathcal{H}_B(\xi) \right)$$

eigenspaces of  $\mathcal{Z}$

• Typical Entropy of Sub-Algebra  $\mathcal{P}(S[\mathcal{O}_A])$

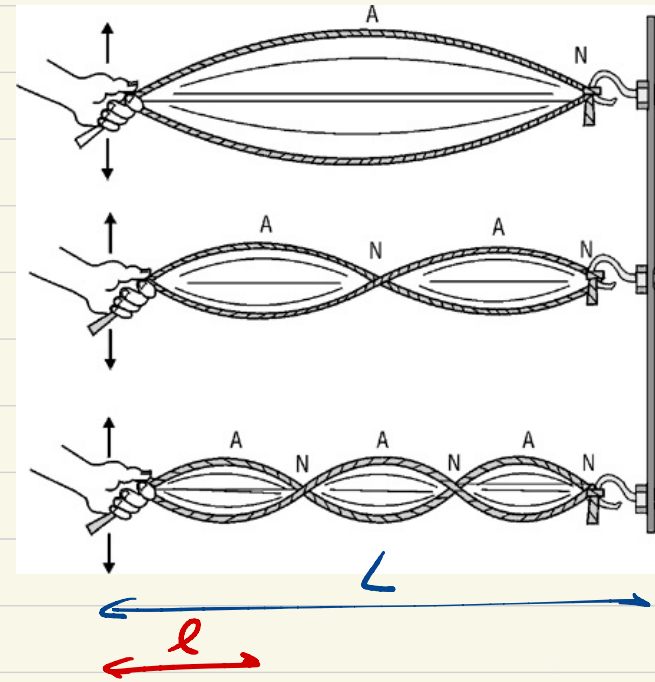
[EB-Satz, 2019]  
[EB-Donà, 2019]

$$\langle S[\mathcal{O}_A] \rangle = \sum_{\xi} \frac{d_{\xi}}{d} \left( \langle S_{\xi} \rangle + \Psi(d+1) - \Psi(d_{\xi}+1) \right)$$



■ Example: Electromagnetic Field & Black Body Radiation

- EM field in a box of volume  $L^3$
- Pure state  $|\psi_E\rangle$  of energy  $E$
- Measurement: Antenna of length  $l$   
 $\hookrightarrow$  sub algebra  $A_A$



- Typical Entanglement Entropy

$$\langle S[A_A] \rangle_E \approx \frac{4\sqrt{\pi}}{3(15)^{1/4}} \left( \frac{\epsilon_A l}{\pi c} \right)^{3/4}$$

$\hookrightarrow$  Black Body Entropy

- Temperature  $kT = \left( \frac{\partial}{\partial \epsilon_A} \langle S \rangle_E \right)^{-1}$

[EB - Donà - Muino, 2022 to appear]

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# ■ Typical Black Hole Entanglement Entropy

- Evaporating Black Hole

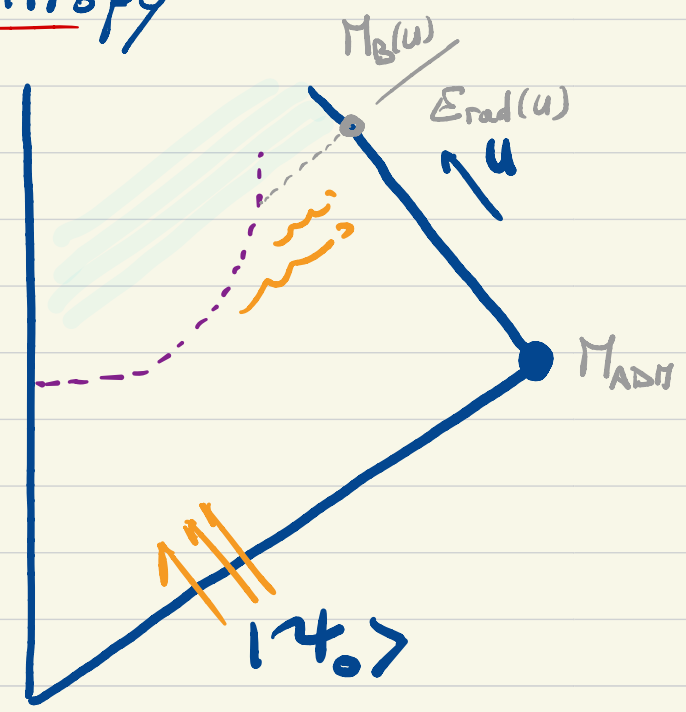
$$M_{\text{ADM}} = M_{\text{B}}(u) + E_{\text{rad}}(u)$$

- Assume pure state for Grav + Matt

$|\Psi_{M_{\text{ADM}}}\rangle$  at fixed  $M_{\text{ADM}}$

- Algebra of Observables  $\mathfrak{g}_{\text{r}}$ ,  $\varphi$

- Sub-algebra:  $\mathfrak{g}_{\text{r}}$  only. Center  $\mathcal{Z}$ :



$\hat{\mathcal{Z}}:$ 

$$\left\{ \begin{array}{l} \hat{\Pi}_{\text{B}}(u) \\ \hat{E}_{\text{rad}}(u) - M_{\text{ADM}} \end{array} \right.$$

## ■ Average Entanglement Entropy of Sub-Algebra

- use  $\dim \mathcal{H}_{\text{BH}}(M_{\text{B}}) \approx \int \mathcal{Z}(\beta) e^{+\beta M_{\text{B}}} \frac{d\beta}{2\pi i} \approx e^{+4\pi \frac{G M_{\text{B}}^2}{\hbar c}}$

Gibbons-Hawking '77  
 Brown-York '93  
 Sen '12  
 EB-Haggard '18

- $\langle S[\text{O}_A] \rangle$  formula in terms of  $\dim$  (EB-Dona '19)

$$\Rightarrow \langle S_{\text{BH}} \rangle \approx 4\pi \frac{G M_{\text{B}}(u)}{\hbar c} = \frac{A_{\text{BH}}(u)}{4G\hbar/c^3}$$

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