

# On Gravitational Dynamics

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NFL : Newton's First Law

Uniform Motion  $\equiv$  No Motion

Absence of Force

- Space is homogeneous

$$x \leftrightarrow y$$

- Time is homogeneous

why not  $x \leftrightarrow t$ ?

Match  $D : x \leftrightarrow ct$

$c$  : Universal constant velocity

- New Mechanics  $\rightarrow$   $SR$
- Spacetime

NFL : No Force! Motion is geodesic of homogeneous spacetime

$$\text{NSL} : m \vec{\ddot{x}} = \vec{F}$$

Galileo's Experiment :

$$\text{Gravity} : m_i \vec{\ddot{x}} = m_g \nabla \phi$$

$m_i = m_g$  why?

No Force  $\xrightarrow{\text{Dual}}$

Universal Force

Links to All

Present everywhere

& Always

Universal property    Universal Property

homogeneous

inhomogeneous

Flat Minkowski

curved spacetime

$$R_{abcd} = 0$$

$$R_{abcd} \neq 0$$

$R_{abcd}$  to determine gravitational dynamics

No need for  $PE$  &  $m_i = m_g$ .

Bianchi identity

Wheeler : Boundary of boundary is zero

$$D^2 = 0 = \nabla \times \nabla \phi = \nabla \cdot \nabla \times \bar{A}$$

Trace is vacuous

Tensor Field :  $g_{ab}$

$$A^b R^a{}_{bcd} = A^a{}_{;cd} - A^a{}_{;dc}$$

$$R_{ab[cd;e]} = 0$$

Now trace is non-vacuous.

$$\Rightarrow G^{ab}{}_{;b} = 0$$

$$G_{ab} = KT_{ab} + \Lambda g_{ab}$$

Einstein's equation follows from curvature

$\Lambda$  : New constant

- Constant curvature of homogeneous & isotropic spacetime — Absence of force
- Signature of spacetime being dynamic not fixed

Einstein's equation also follows from

$$\begin{aligned}
 & \delta \int R_{ab} g^{ab} \sqrt{-g} d^4x \\
 &= \int [R_{ab} \sqrt{-g} \delta g^{ab} + R \delta \sqrt{-g} + (\delta R_{ab}) g^{ab} \sqrt{-g}] d^4x \\
 &= \int (R_{ab} - \frac{1}{2} R) \sqrt{-g} d^4x + \int \delta R_{ab} = 0
 \end{aligned}$$

Variation of dynamic part,  $R_{ab}$  makes no contribution.

Higher order in curvatures.

Find an analogue of  $R_{abcd}$  satisfying Bianchi identity.

$$\mathcal{R}_{abcd} = R_{abmn} R_{cd}{}^{mn} + \alpha R_{[a}{}^m R_{b]mcd} + \beta R R_{abcd}$$

Bianchi :  $\mathcal{R}_{ab[cd;e]}$  to yield a divergence free tensor requires

$$\alpha = 4, \beta = 1$$

Guass -Bonnet term.

$$\mathcal{R}^{cd}_{[cd;e]} = \frac{1}{2}\mathcal{R}_{,e} \neq 0$$

$$\mathcal{R}^{cd}_{[cd;e]} - \frac{1}{2}\mathcal{R}_{,e} = -H_{e;c}^c = 0$$

This suggests

$$F_{abcd} = \mathcal{R}_{abcd} - \gamma\mathcal{R}(g_{ac}g_{bd} - g_{ad}g_{bc})$$

$$F^{cd}_{[cd;e]} = H_{e;c}^c = 0$$

This requires

$$\mathcal{R}_{abcd} = Q_{ab}{}^{mn} R_{mncd}$$

$$Q_{cd}{}^{ab} = \delta_{cdc_1d_1 \dots c_nd_n}{}^{aba_1b_1 \dots a_nb_n} R_{a_1b_1}{}^{c_1d_1} \dots R_{a_nb_n}{}^{c_nd_n}.$$

For quadratic GB

$$Q_{abcd} = R_{abcd} - 2R_{a[c}g_{d]b} + 2R_{b[c}g_{d]a} + Rg_{a[c}g_{d]b}.$$

In general

$$F_{abcd} = \mathcal{R}_{abcd} - \frac{n-1}{n(d-1)(d-2)} \mathcal{R}(g_{ac}g_{bd} - g_{ad}g_{bc})$$

$n$  : order of polynomial

$d$  : dimension of spacetime

$$n(F_{ab} - \frac{1}{2}Fg_{ab}) = H_{ab}$$

$$H_{a;b}^b = 0.$$

$$n = 1, F_{ab} = R_{ab}, G_{ab} = H_{ab}$$

**Theorem :** *The second order quasi-linear differential operator as a second rank divergence free tensor in the equation of motion for gravitation could always be derived from the trace of the Bianchi derivative of the fourth rank tensor,  $F_{abcd}$ , which is a homogeneous polynomial in curvatures. The trace of the curvature polynomial is proportional to the corresponding term in the Lovelock action and corresponding to each term in the Lovelock Lagrangian, there exists a fourth rank tensor which is a new characterization of the Lovelock Lagrangian.*

## Lovelock Polynomial

- quasi-linearity
- Metric and Palitini Variation  
→ the same equation.
- Bianchi identity → Equation.

New identification of Lovelock gravity

For  $d > 4$ ,  $H_{ab}$  is as natural as  $G_{ab}$  and hence must be included.

Equation to follow from  $F_{abcd}$  → potential for  $H_{ab}$

GB :

1. One loop (string) correction
2. Higher D
3. High Energy

CG  $\rightarrow$  GB / Lovelock  $\rightarrow$  QG

GB : Intermediatory Limit

QG / High Energy  $\rightarrow$  Higher D?

## **Higher D**

- Isometric Flat Space Embedding
- Self interaction iteration

First :  $(\partial)^2$

Second :  $(\partial)^4 : R_{abcd}^2$

$\Rightarrow$  GB  $\rightarrow$  Relevant only in  $D > 4$ .

Higher D required.

How to see higher D?

Matter Probes confirmed to 4 - D

Only gravity in higher D

Purely gravitational experiment?

Signature in sub  $mm$  / cosmology?

**Search is on.**